

Exploring Habitability and Radio Emission in Proxima b's Space Weather Environment

Luis Peña Moñino

Supervisor: Miguel Pérez Torres

lpm@iaa.es



CSIC

I acknowledge funding through grant [PRE2020-095421](#), funded by MCIU/AEI/10.13039/501100011033 and by FSE Investing in your future. I also acknowledge financial support through the Severo Ochoa grant [CEX2021-001131-S](#) and the National grant [PID2020-117404GB-C21](#), funded by MCIU/AEI/ 10.13039/501100011033.

M-dwarfs and radio emission



Stellar activity

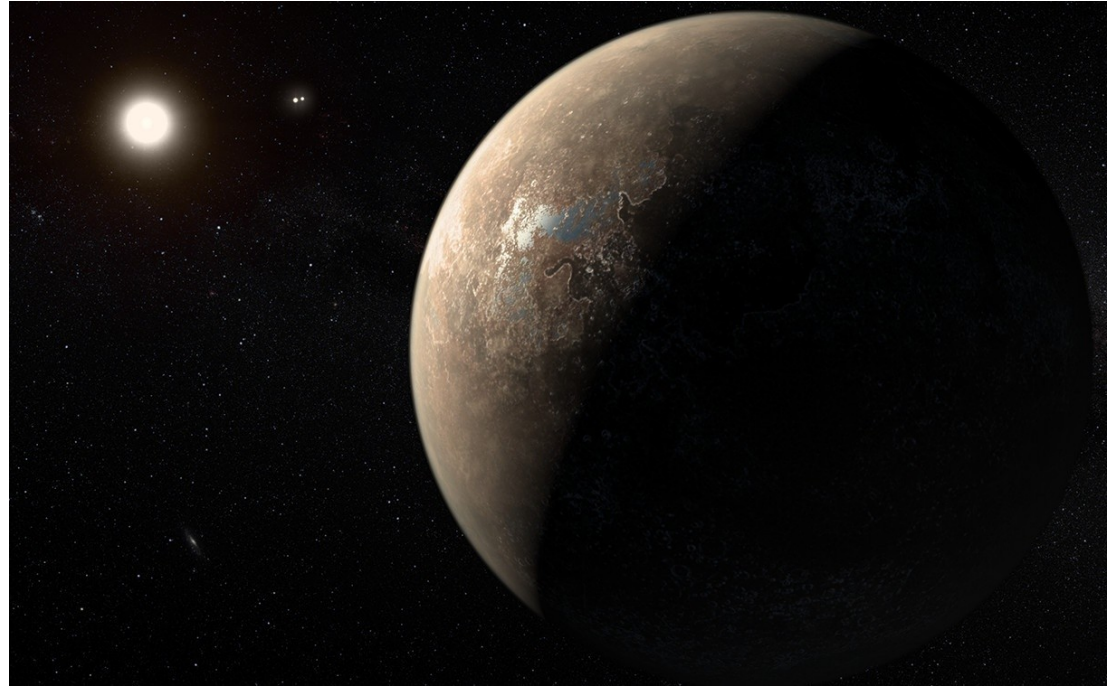
Coronal mass ejections, flares
Magnetic field
Impact on planet habitability

Planetary magnetic field

Shield
Habitability indicator
Informs on planet interior

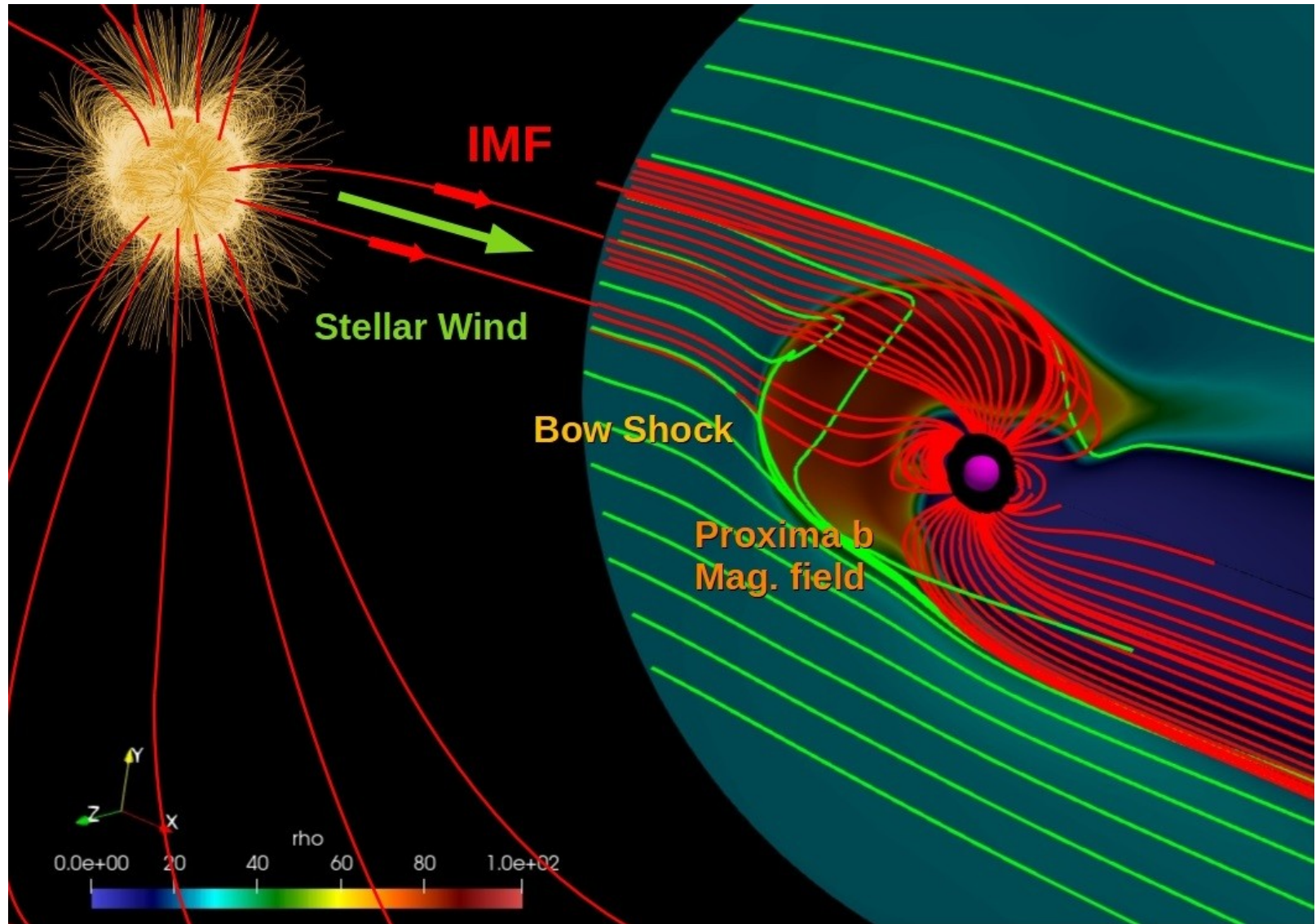
PLUTO simulations of the space weather in Proxima b

- Safe for life?
- Direct emission from the magnetosphere?



Peña-Moñino et al (2024), A&A 688, A138.
DOI: [10.1051/0004-6361/202349042](https://doi.org/10.1051/0004-6361/202349042)

PLUTO simulations of the space weather in Proxima b



PLUTO simulations of the space weather in Proxima b

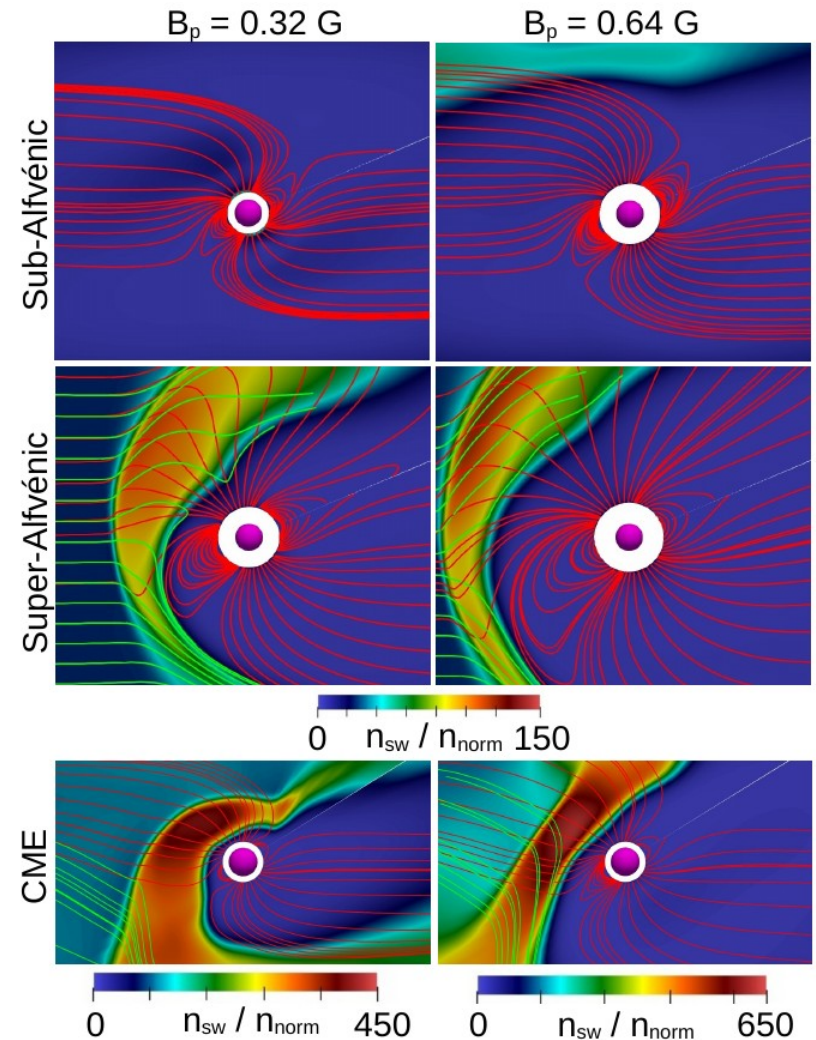
WIND PARAMETERS

Parameter	Calm space weather		Extreme space weather
	Sub-Alfvénic scenario	Super-Alfvénic scenario	CME scenario
n_{sw} [cm^{-3}]	50	50	250
$ v_{\text{sw}} $ [10^7 cm s^{-1}]	5	10	25
B_{IMF} [mG]	3.2	1.6	16
B_{Proxima} [G]	1200	600	600
$B_{\text{Proxima b}}$ [G]	0.16–1.28	0.16–1.28	0.16–0.64
i	$[0^\circ - 90^\circ]$	$[0^\circ - 90^\circ]$	$[0^\circ - 45^\circ]$

PLUTO simulations of the space weather in Proxima b

PLANET PARAMETERS

- Inclination: 0 to 90°
- B_{planet} : 0.16 to 1.28 G



PLUTO simulations of the space weather in Proxima b

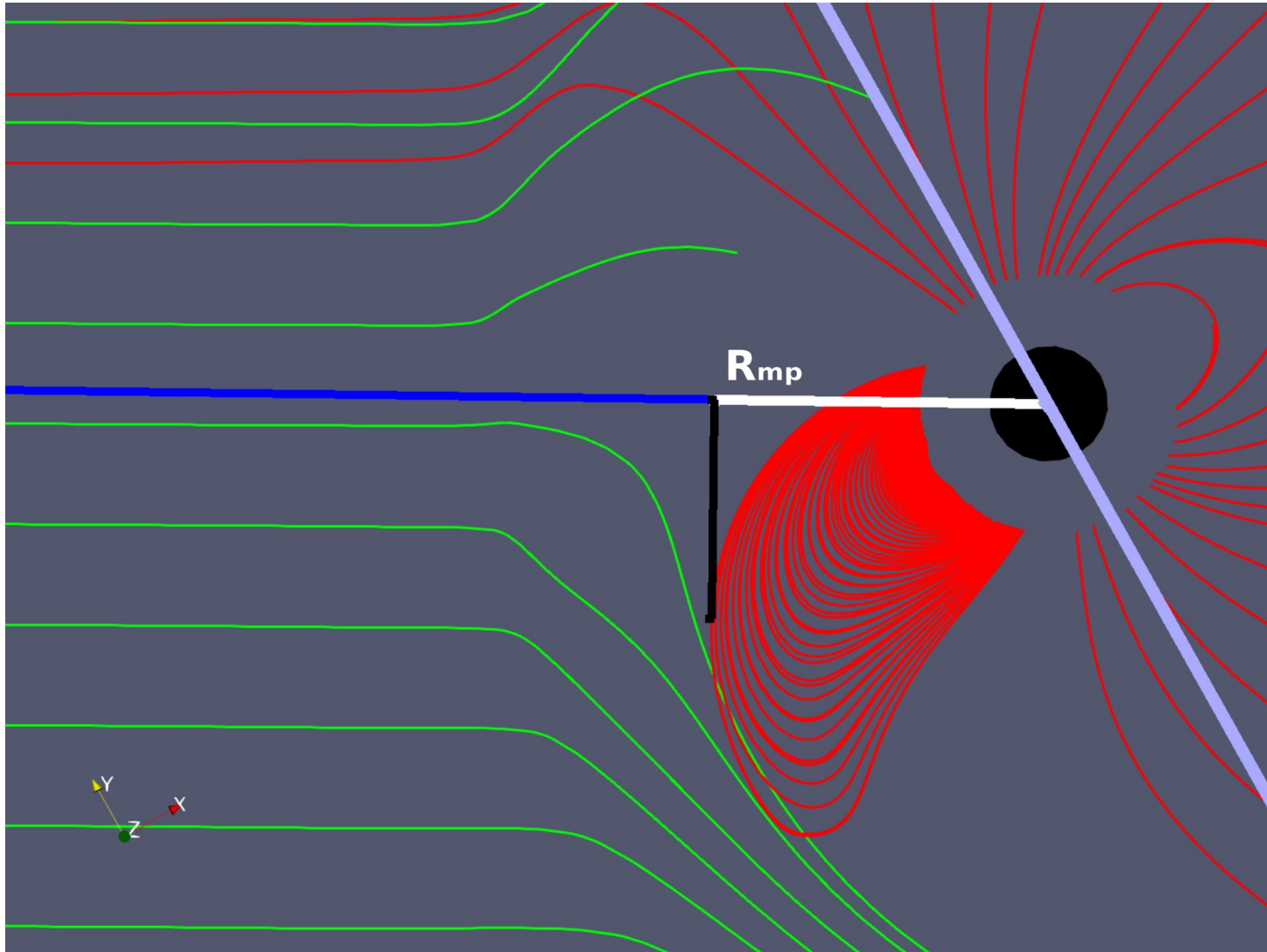
R_{mp} to infer habitability

If $R_{\text{mp}} < R_p$: surface sterilized

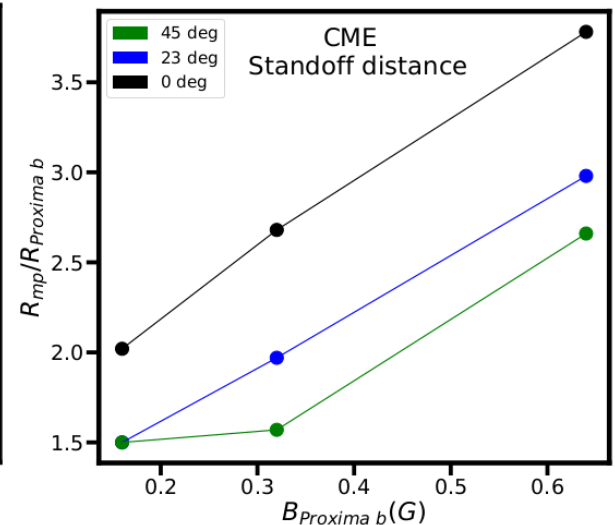
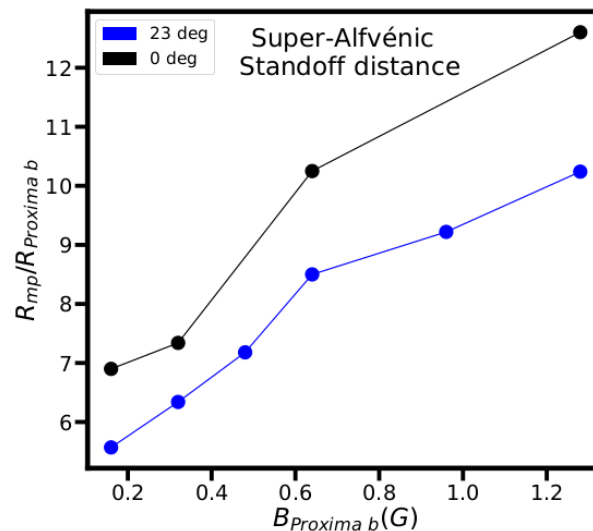
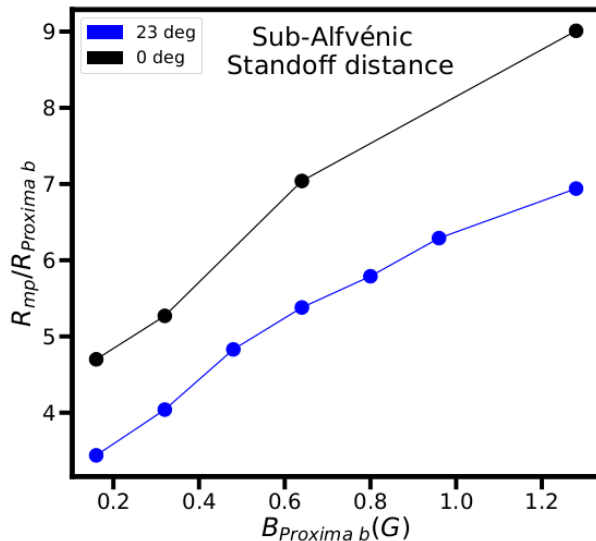
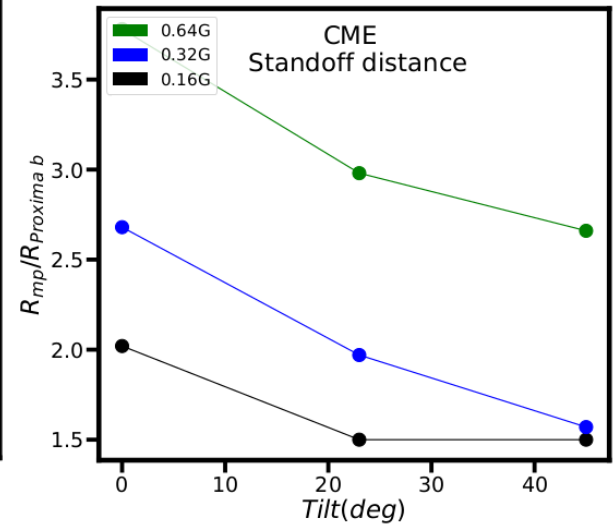
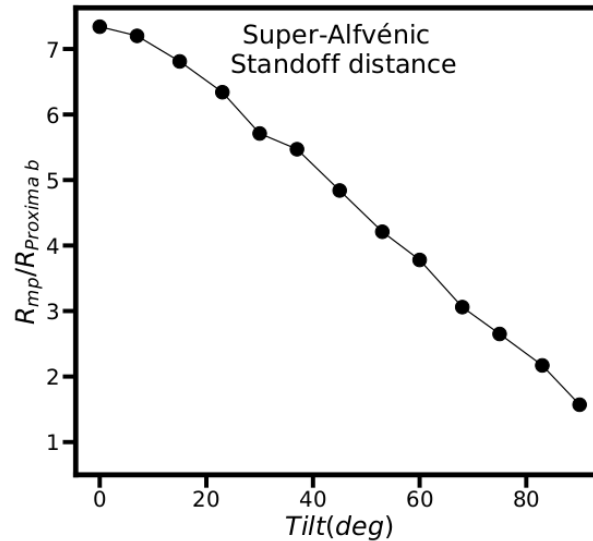
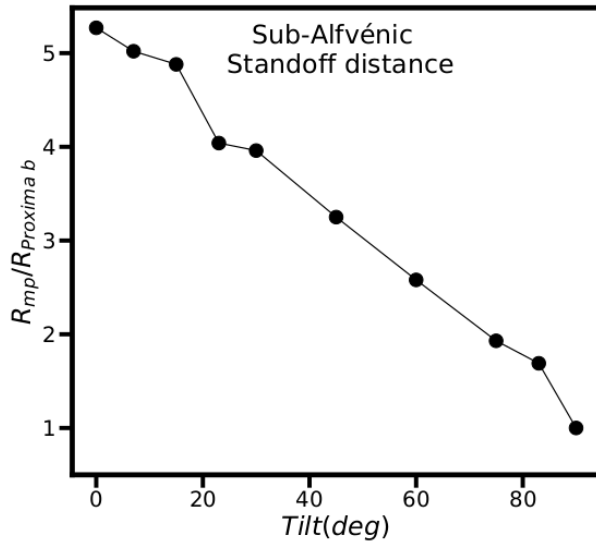
$$\frac{R_{\text{mp}}}{R_p} = \left[\frac{\alpha \mathcal{M}_p^2 / \pi}{m_p n_{\text{sw}} v_{\text{sw}}^2 + \frac{B_{\text{IMF}}^2}{4\pi} + \frac{2m_p n_{\text{sw}} c_{\text{sw}}^2}{\gamma} - m_p n_{\text{bs}} v_{\text{th,msp}}^2} \right]^{1/6}$$

Varella et al. 2022

PLUTO simulations of the space weather in Proxima b

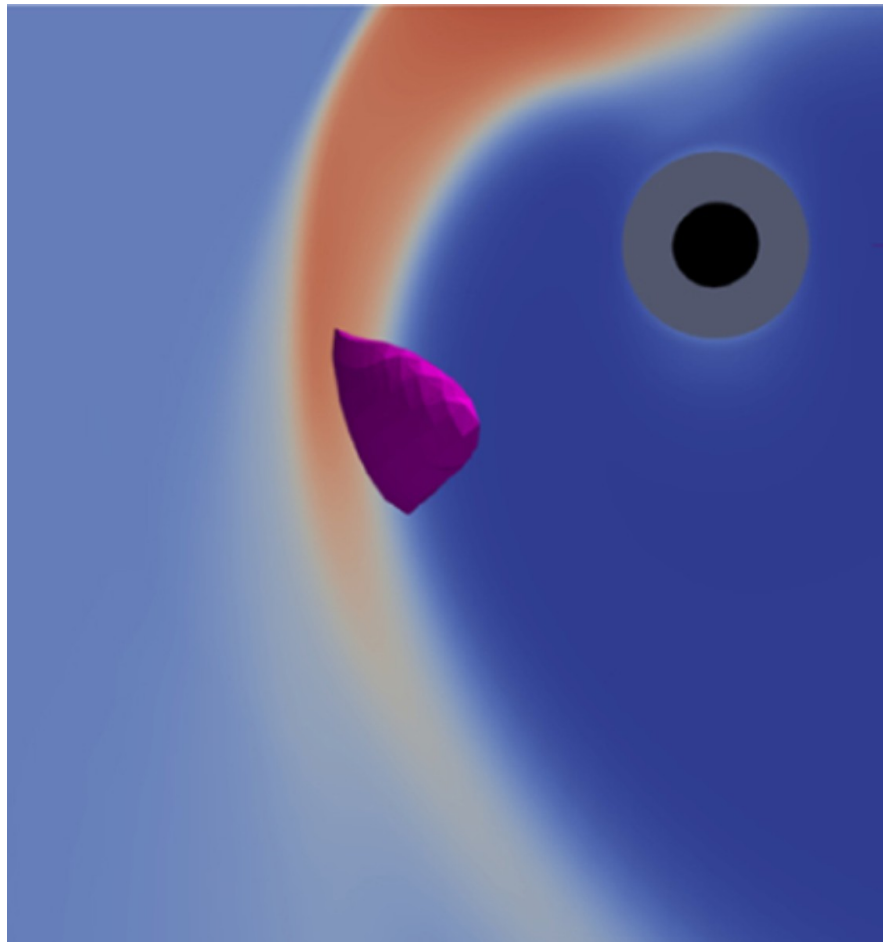


PLUTO simulations of the space weather in Proxima b



PLUTO simulations of the space weather in Proxima b

RADIO EMISSION



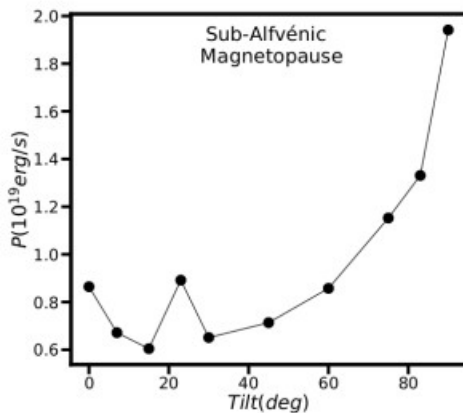
Peña-Moñino et al 2024

PLUTO simulations of the space weather in Proxima b

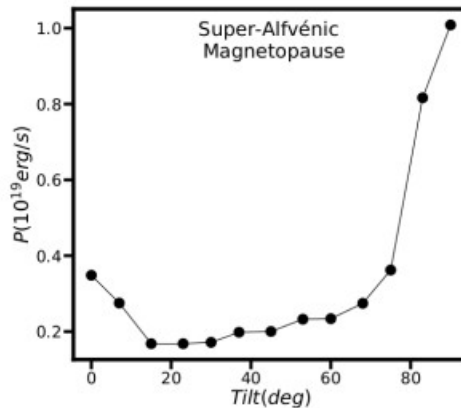
RADIO EMISSION

CALM WEATHER

Sub-Alfvénic

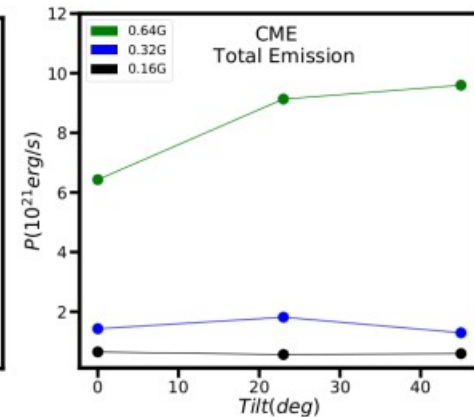
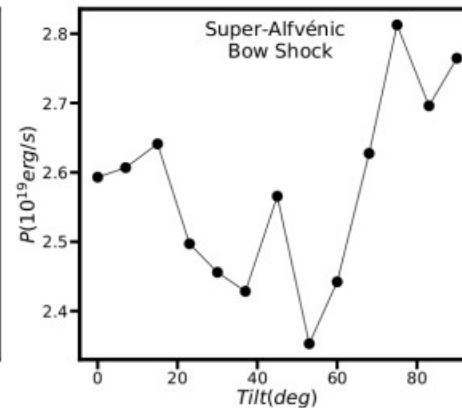


Super-Alfvénic



EXTREME

CME



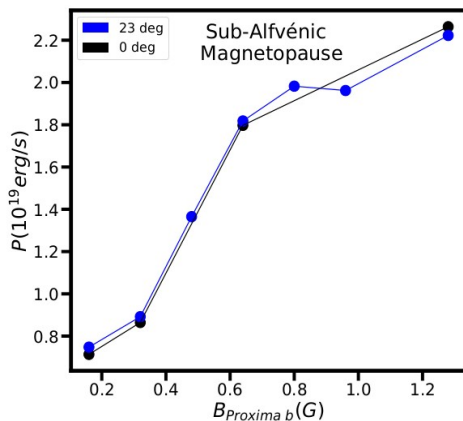
Peña-Moñino et al 2024

PLUTO simulations of the space weather in Proxima b

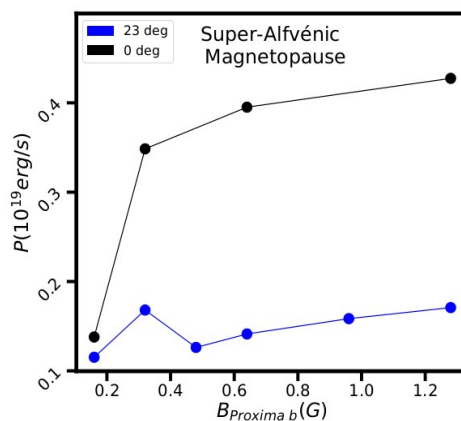
RADIO EMISSION

CALM WEATHER

Sub-Alfvénic

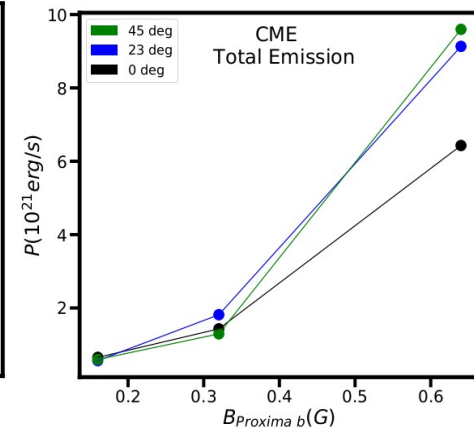
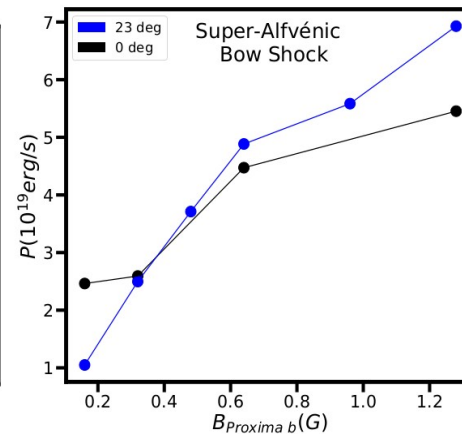


Super-Alfvénic



EXTREME

CME



Peña-Moñino et al 2024

PLUTO simulations of the space weather in Proxima b

- R_{mp} decreases with tilt and increases with B_p . Larger in the super-Alfvénic regime: Magnetic erosion stronger kinematic compression.
- Proxima b **shielded** for most cases.
- **Emission**: bow shock-dominated emission. $10^{18} - 10^{22}$ egs/s

THANK YOU

Contact me at
lpn@iaa.es