

A map of Europe with various radio astronomy stations marked by orange dots. Lines connect these stations, forming a network. A red arrow points from the text area to a cluster of stations in Central Europe. The stations labeled include Birr, Chilbolton, Dutch stations, Norderstedt, Potsdam, Borówiec, Łazy, Tautenburg, Unterweilenbach, Nançay, Effelsberg, Jülich, Onsala, and Irbene. A yellow line also connects the text area to the Nançay station.

Ground-based perspective of PRE radioastronomy, from historical observations to modern giant facilities

A photograph of a radio telescope array, showing numerous large, triangular metal structures arranged in a field. The structures are made of metal poles and cross-bracing, forming a grid-like pattern. The background shows a line of trees under a clear sky.

Philippe Zarka

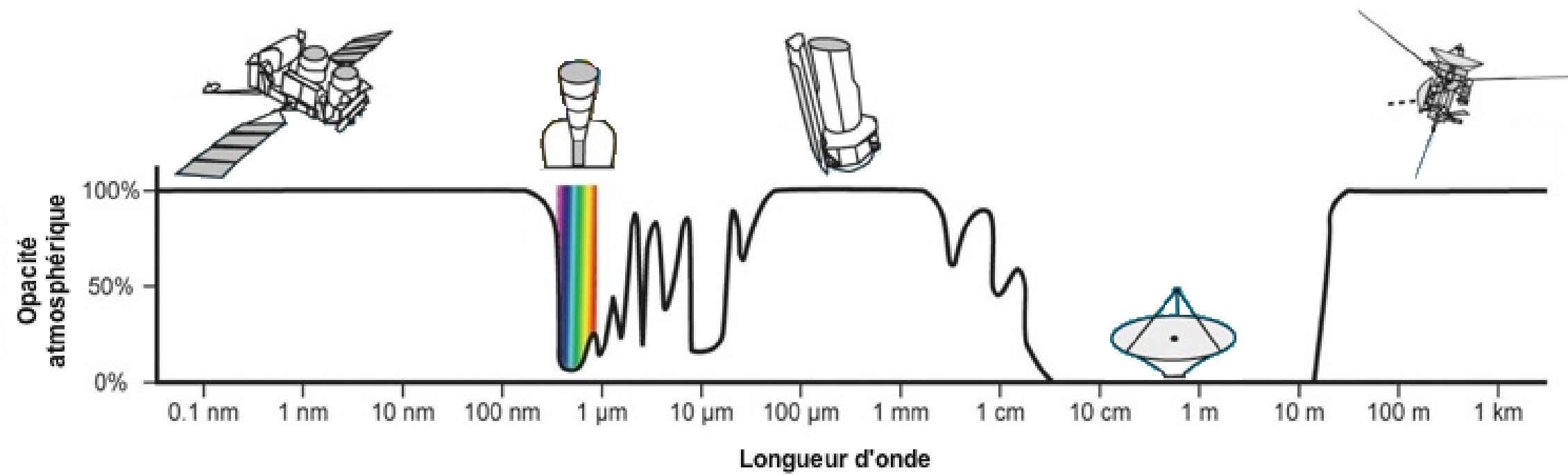
LIRA & ORN, CNRS - Observatoire de Paris - PSL

- Heroic times
 - Jupiter, Sun
- Single dishes and phased arrays: t-f studies
 - Jupiter (<1990, 1990-2000, 2000-2010, 2010-2020, 2020+)
 - Sun, Stars, Exoplanets
- Interferometers: imaging studies
 - Jupiter, Sun, Stars, Exoplanets
- Dynamic spectra (t-f) from imaging data
 - Stars, Exoplanets
- Prospects

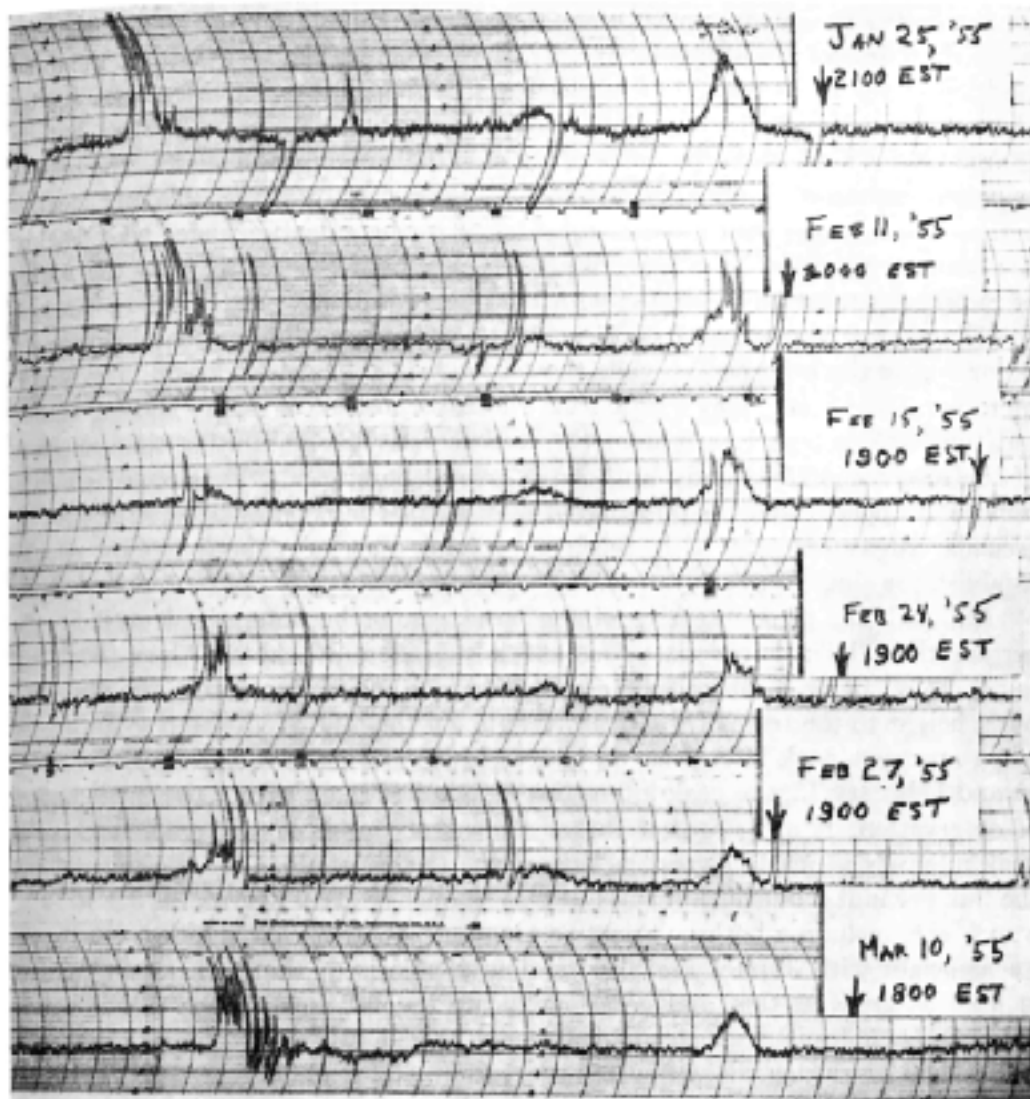
A sketchy, biased review ...

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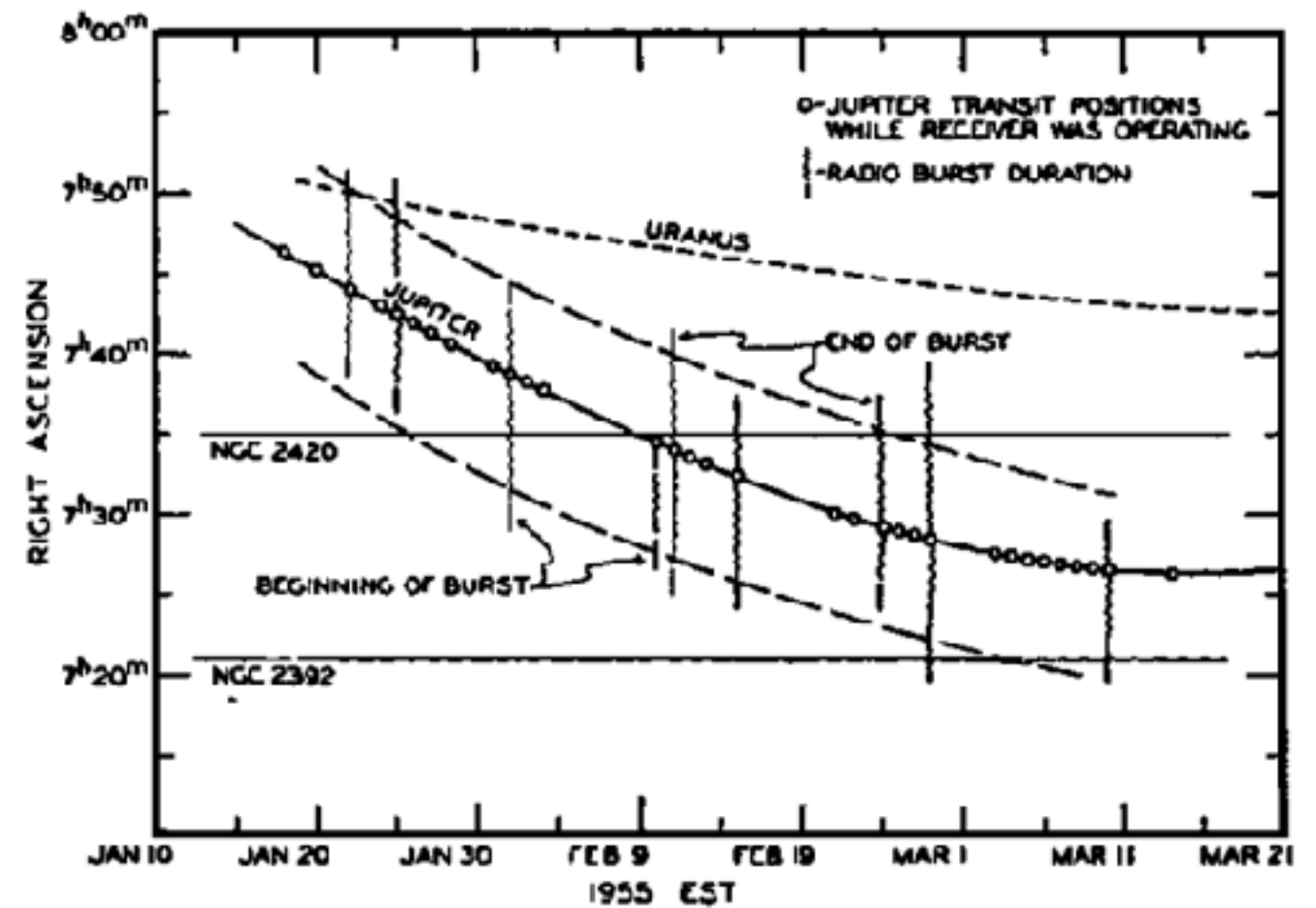
Heroic times



Jupiter



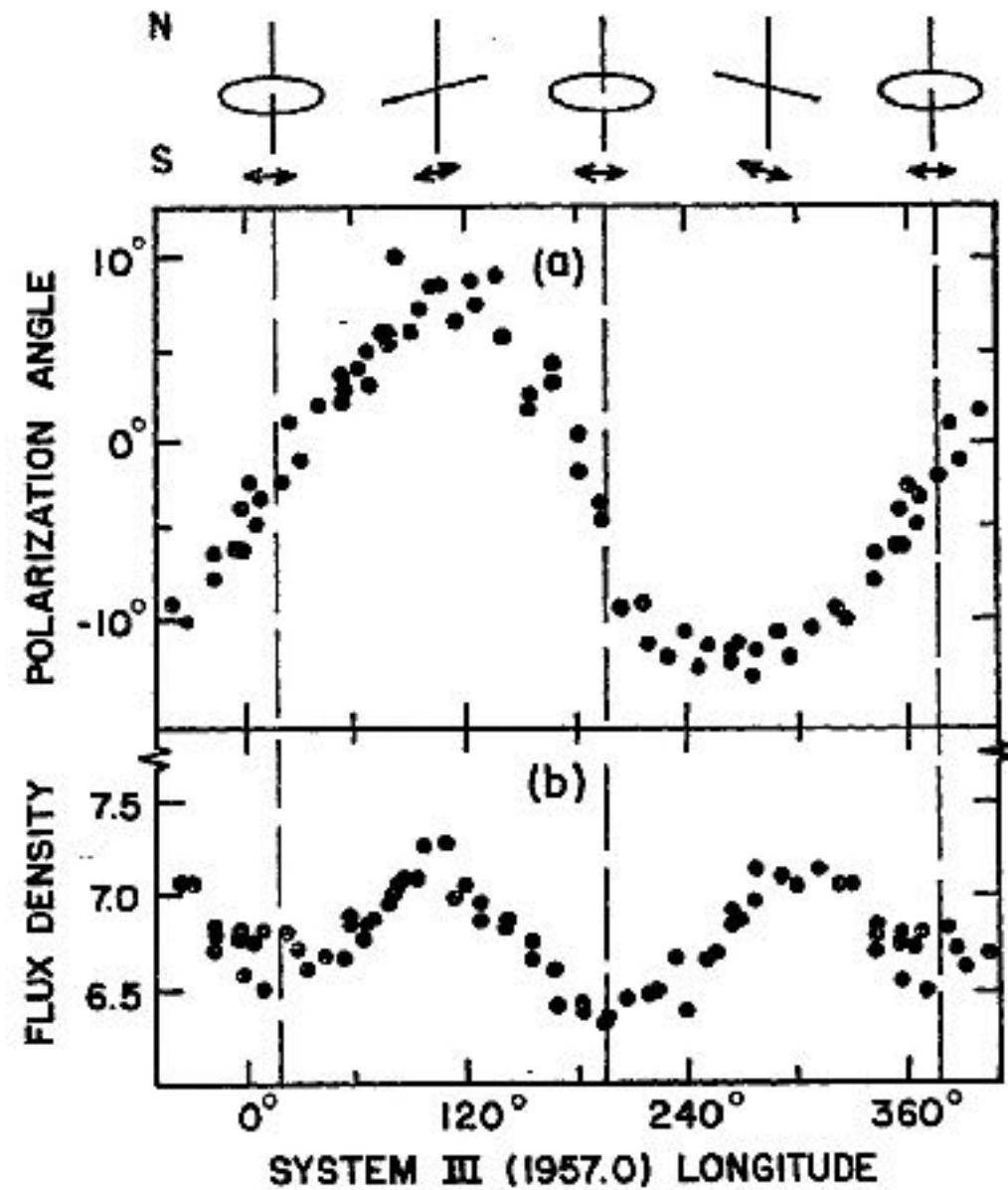
Mills Cross of Carnegie Institution, Washington, 22.2 MHz



[Burke & Franklin, 1955; Franklin & Burke, 1956]

→ electron cyclotron in 10 G B-field

Jupiter

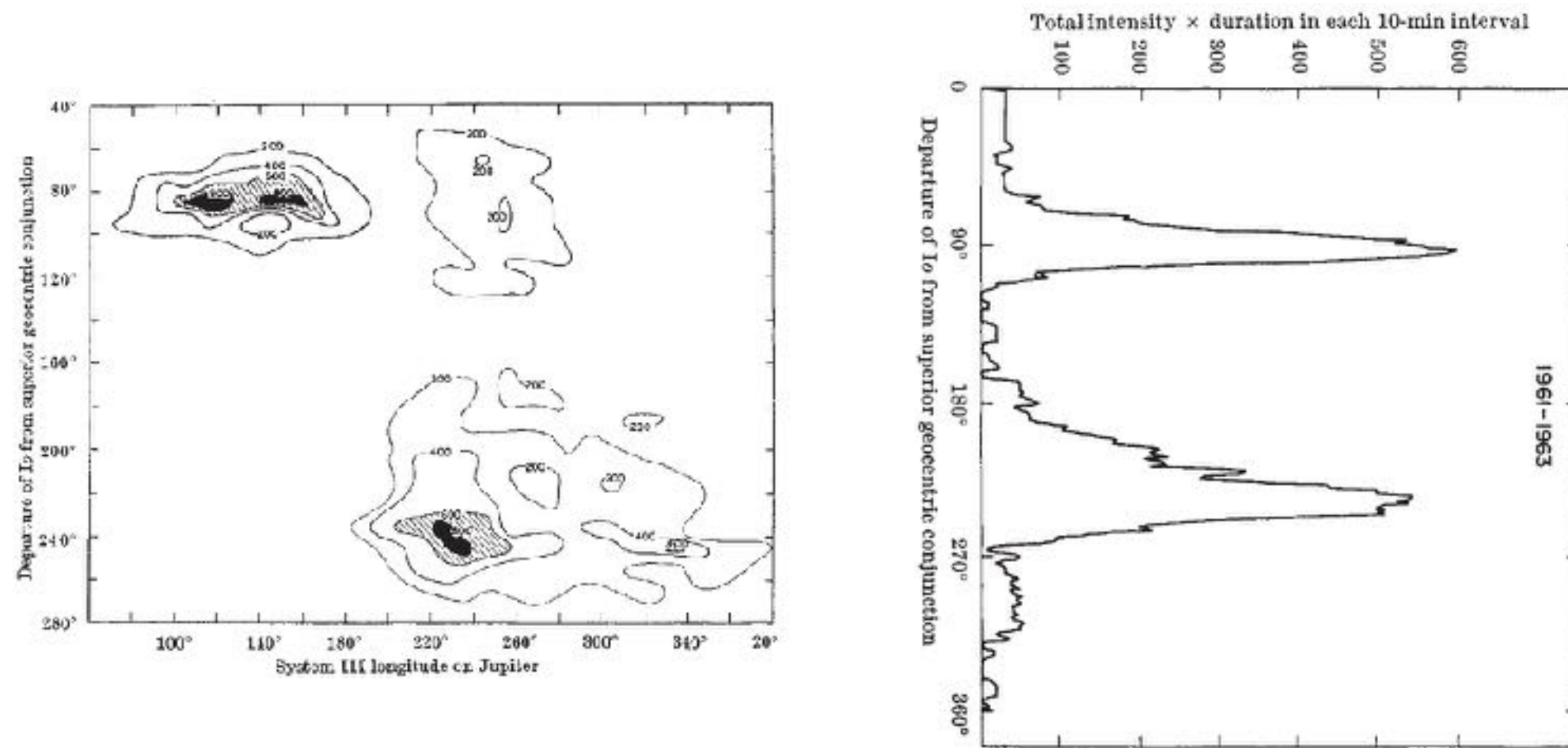


1-3 GHz

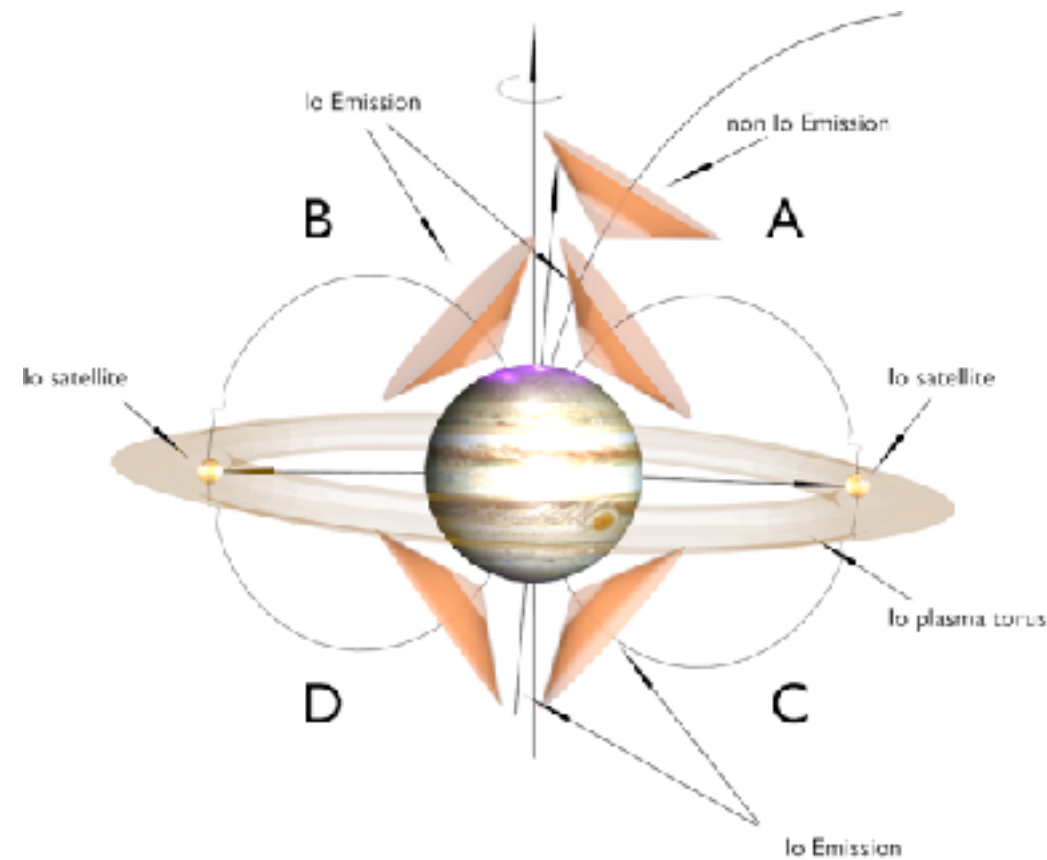
[Sloanaker, 1959; Roberts & Stanley, 1959; Roberts & Komesaroff, 1964]

→ Jovian B dipole tilted by 10°

Jupiter

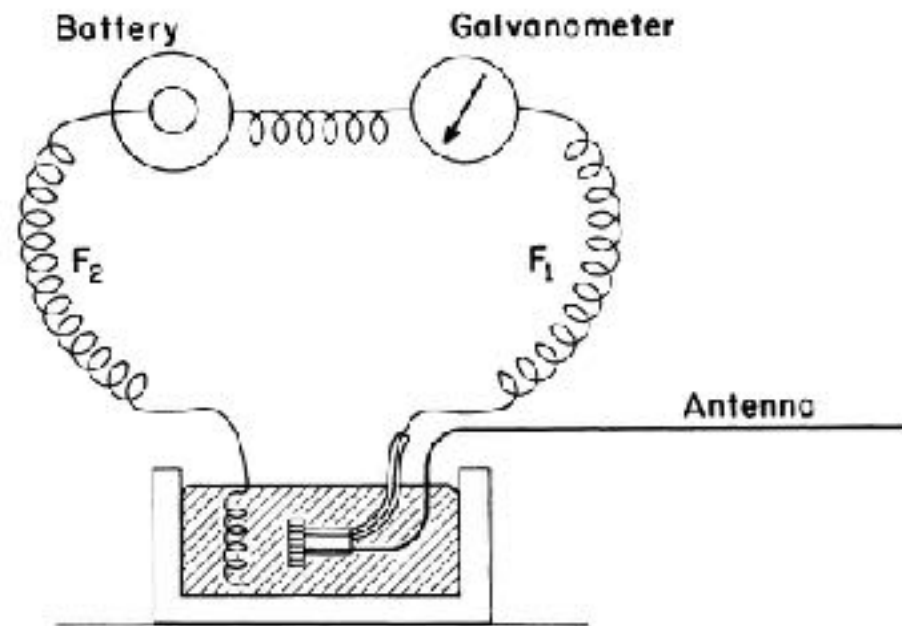


[Bigg, 1964]

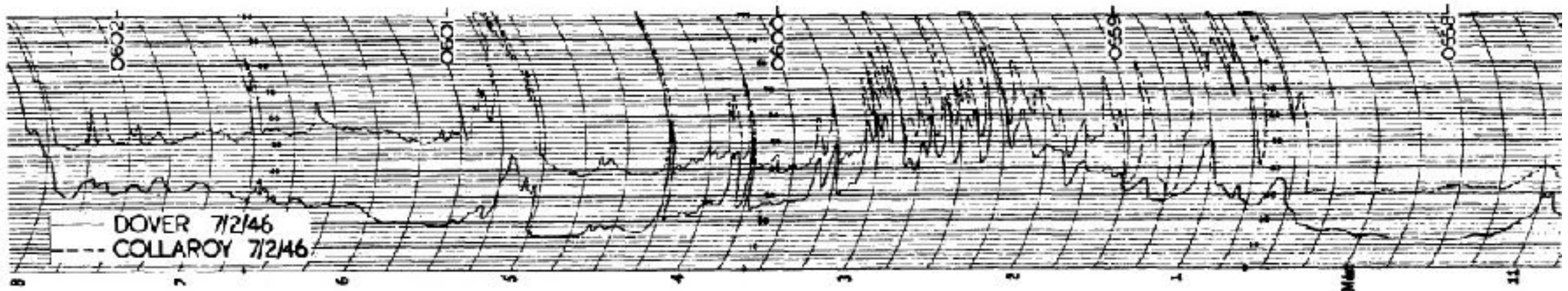


[Marques et al., 2017]

Sun



[Nordmann, 1902 ; Deslandres & Décombe, 1902]



[McCready et al., 1947]

→ **Solar emission at 50-75 MHz** [Hey, 1946]

& 200 MHz

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Single dishes & phased arrays: t-f studies



NDA



UTR-2

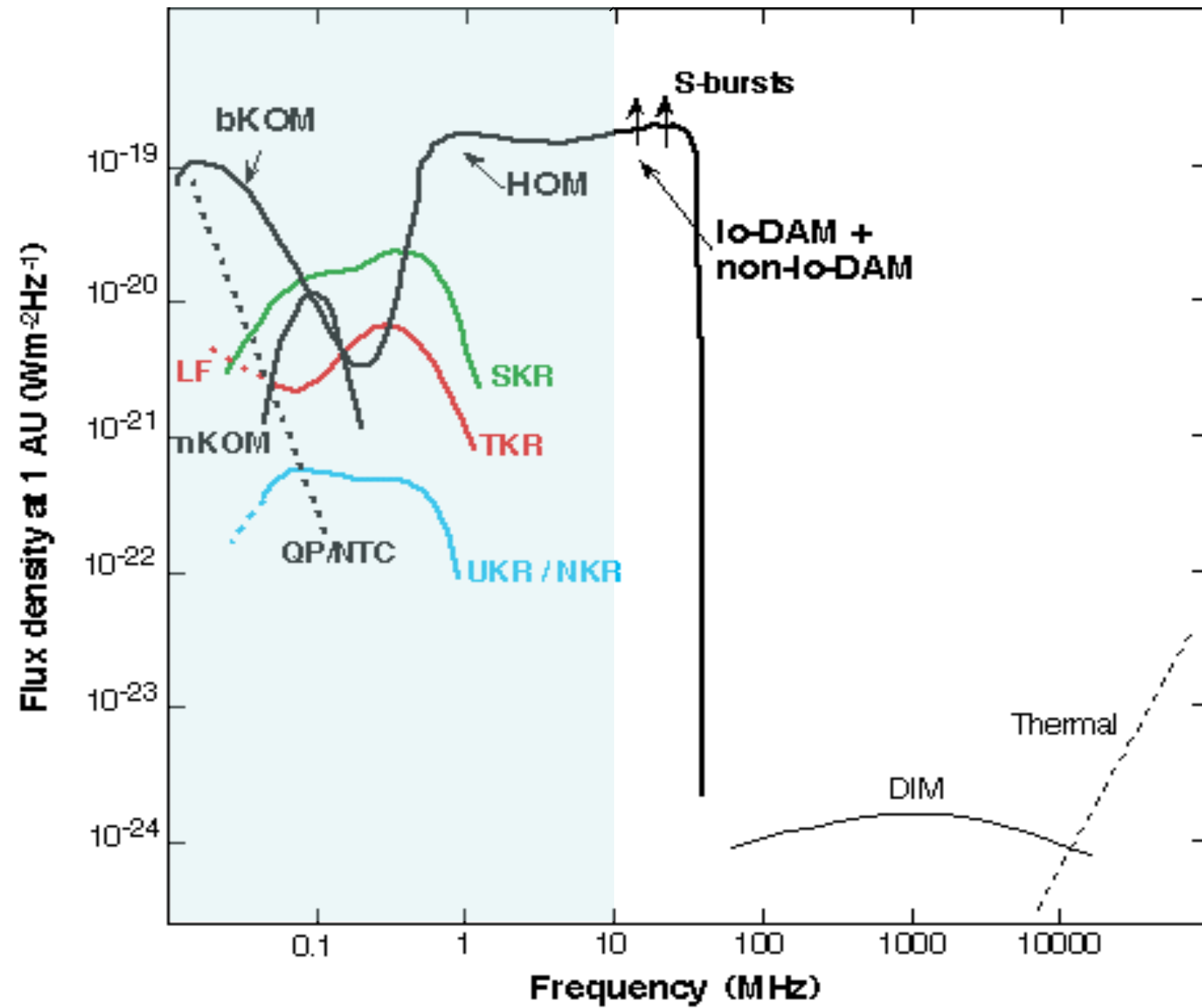


LWA



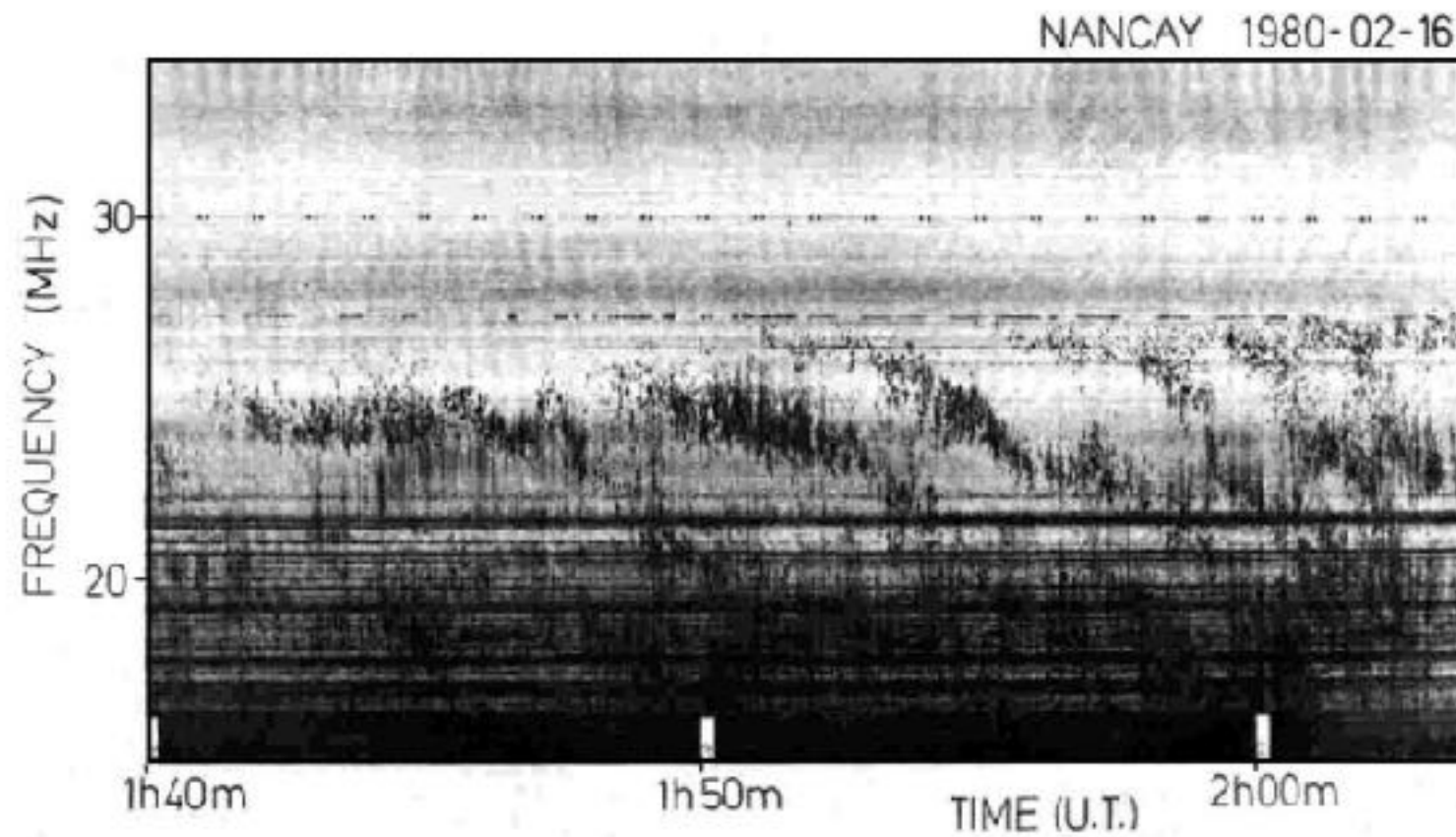
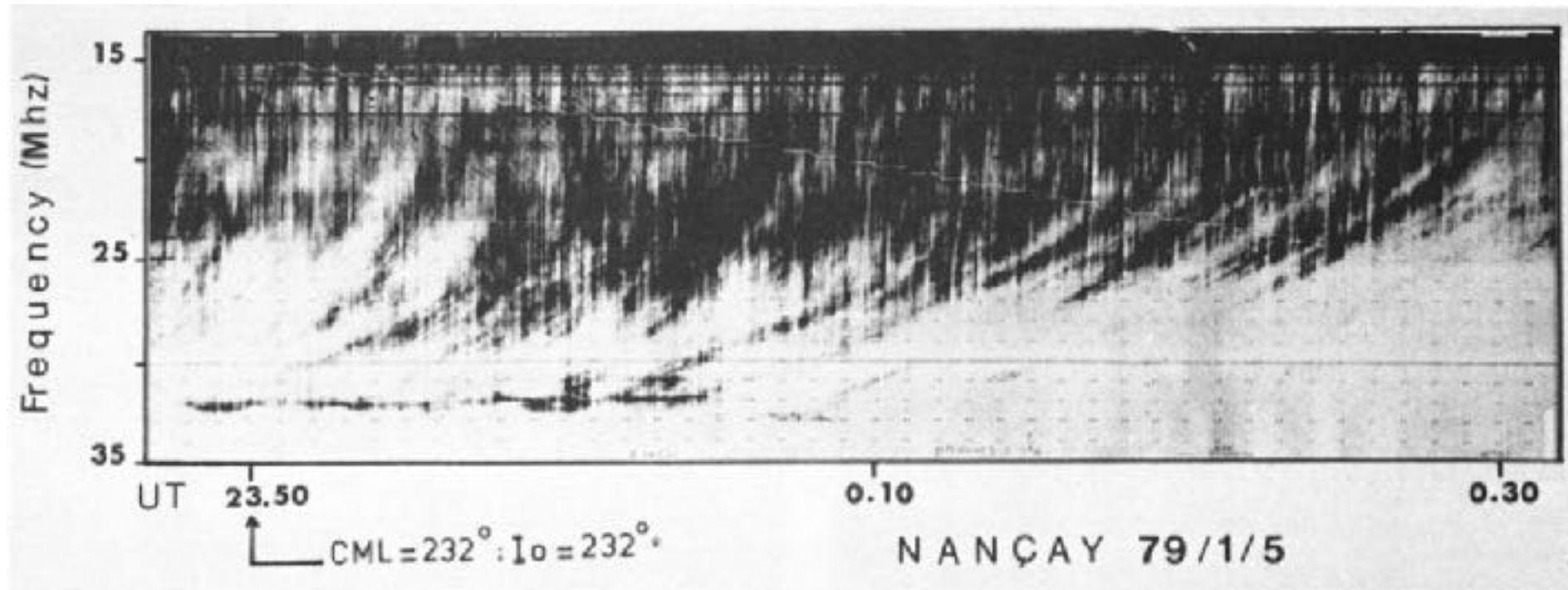
Arecibo

Jupiter



[Zarka, 1998, 2000]

Jupiter: <1990



→ analog records

Jupiter: <1990

- Catalogs & occurrence statistics [Leblanc et al., A&Supp 1981, 1983, 1989, 1990, 1993]
- Arcs morphology ? [Lecacheux et al. 1981, Gurnett & Goertz, 1982]
- Scintillation (ionosphere, IPS) → source location [Genova et al., 1981]
- S-bursts occurrence & interpretation as upgoing e- beams [Ellis, 1980; Leblanc & Genova, 1981]

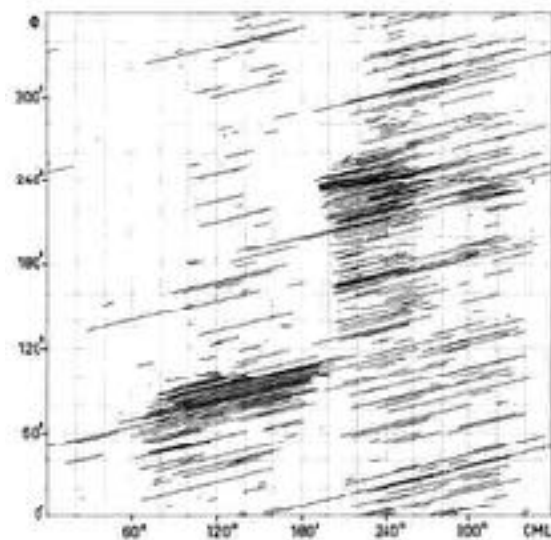
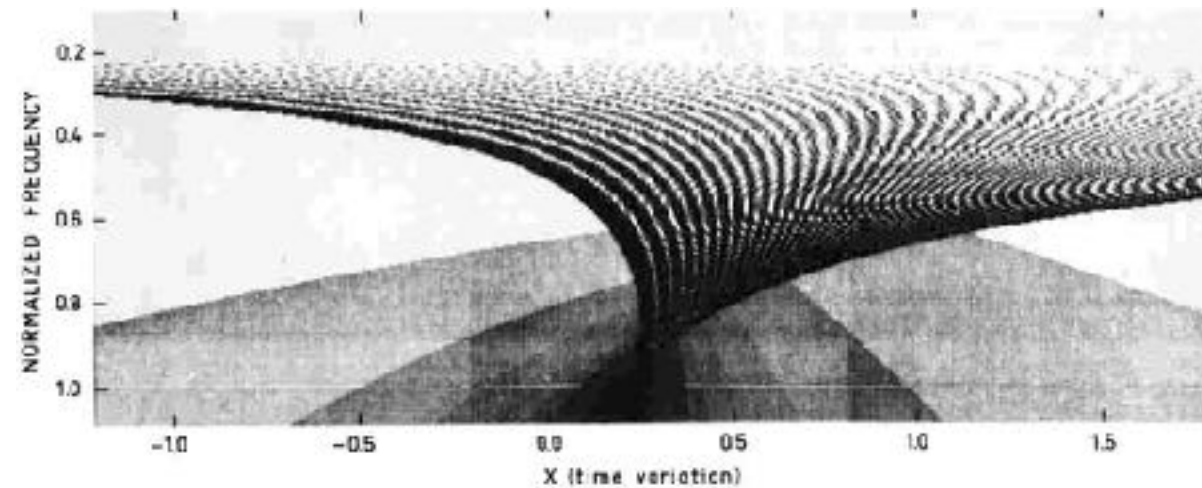
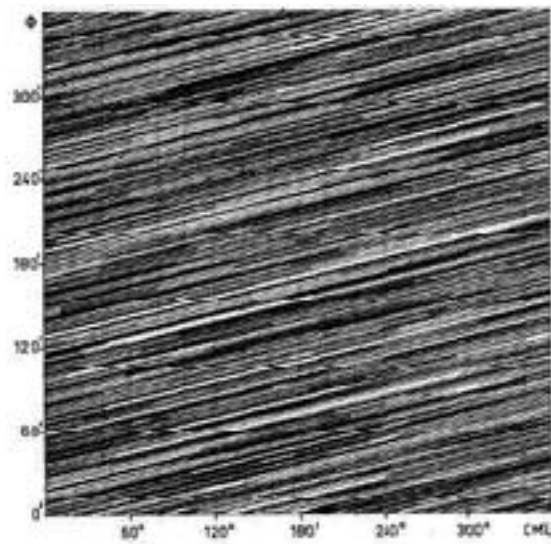
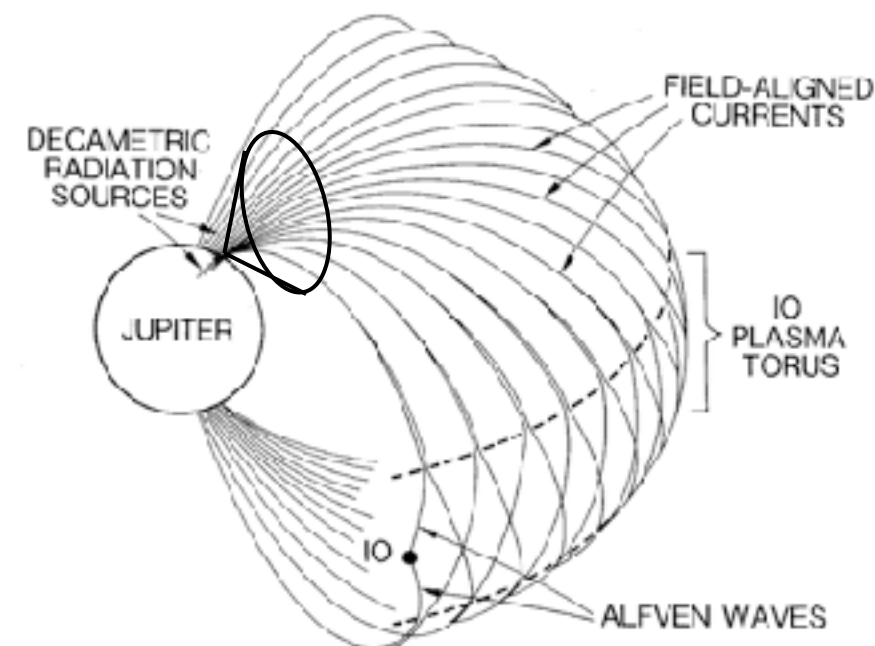
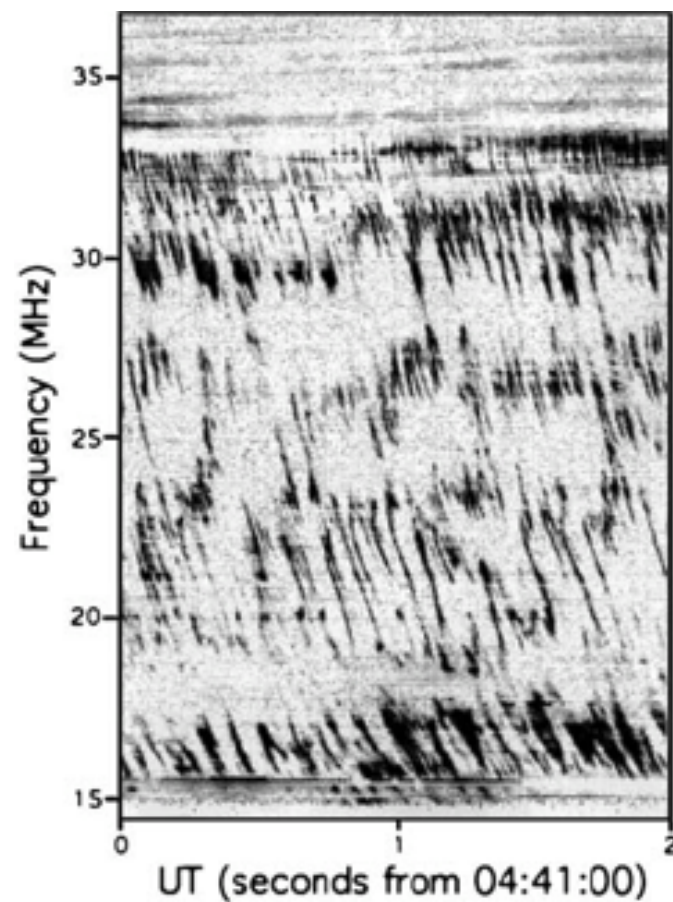
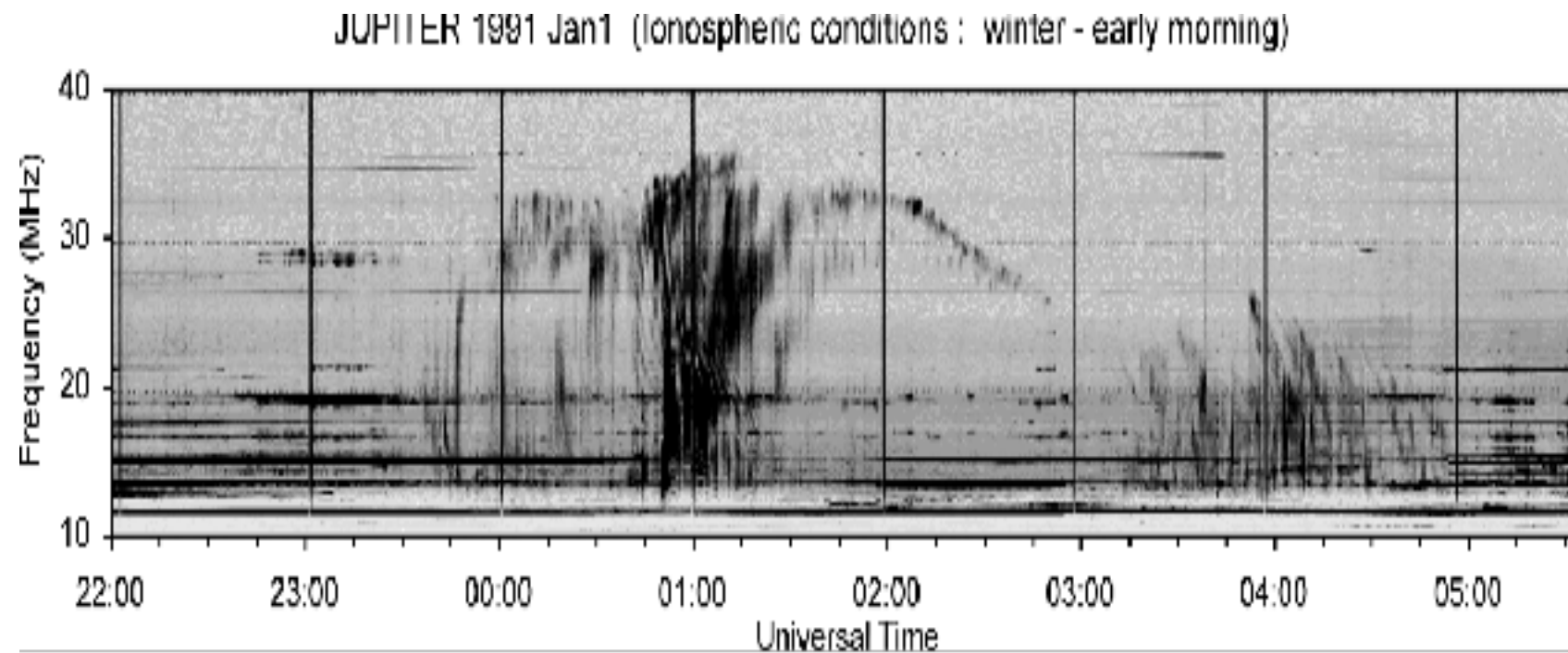


FIGURE 6. — The CML and Io-phase diagram for the period of January 1978 to December 1979. a) the observation tracks; b) the emission tracks.



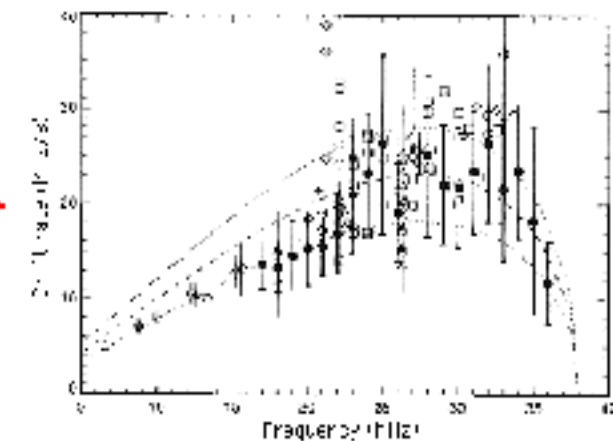
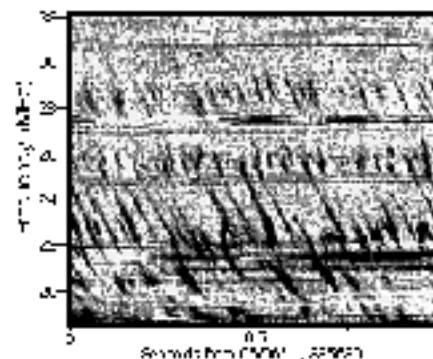
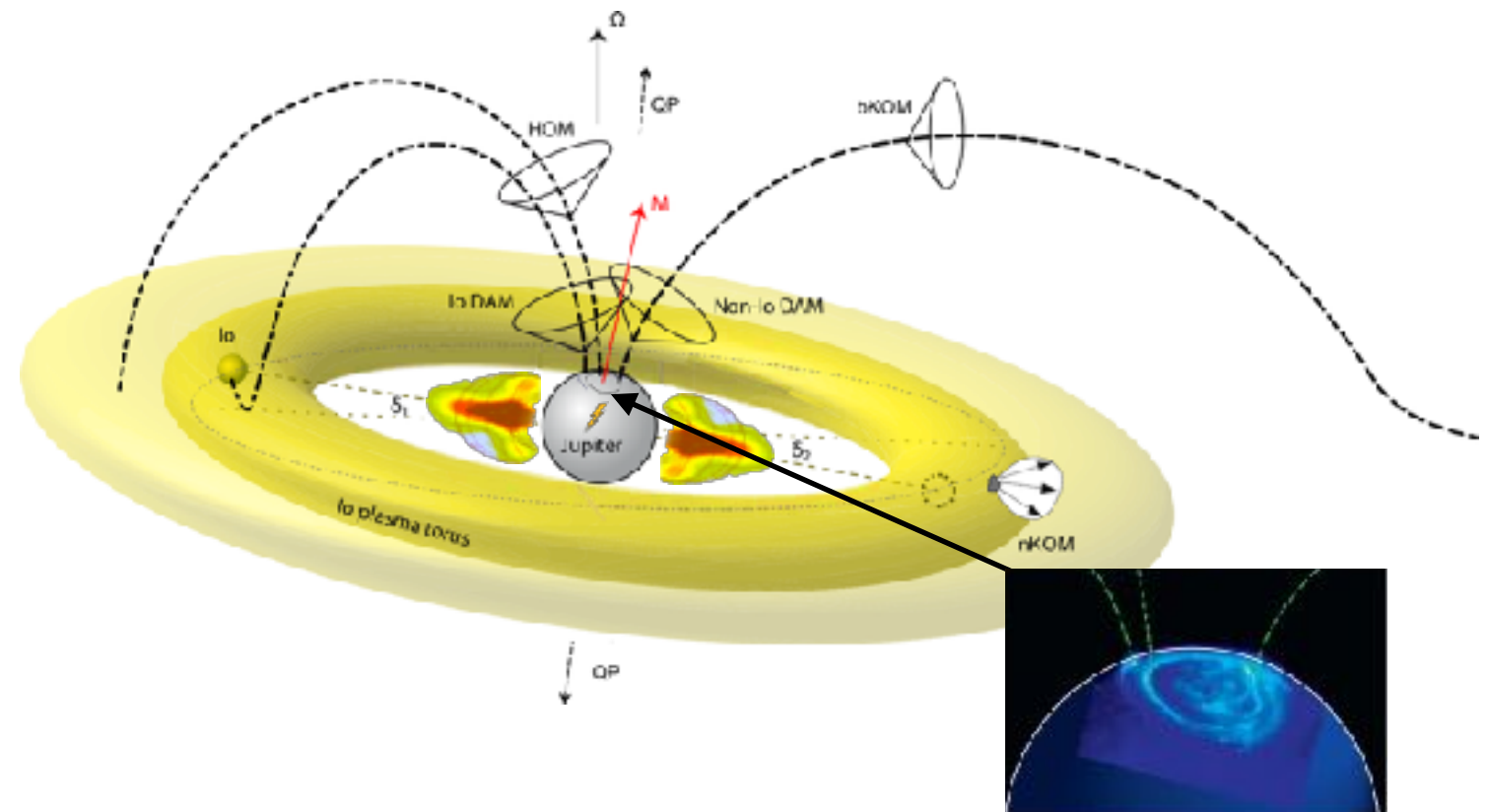
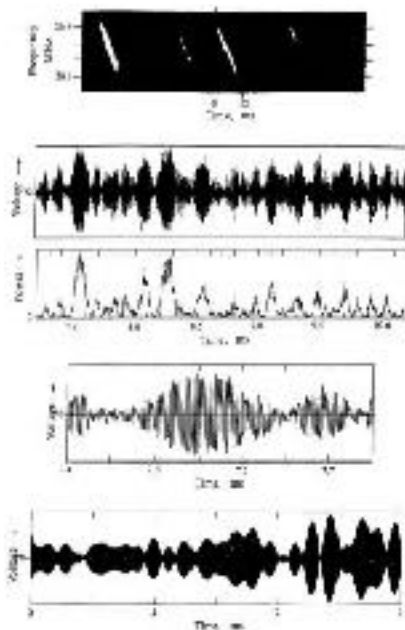
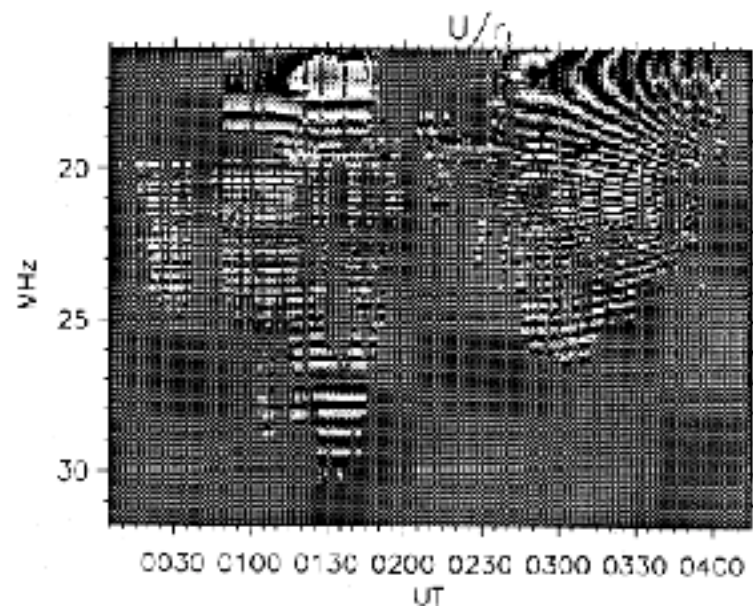
Jupiter: 1990-2000



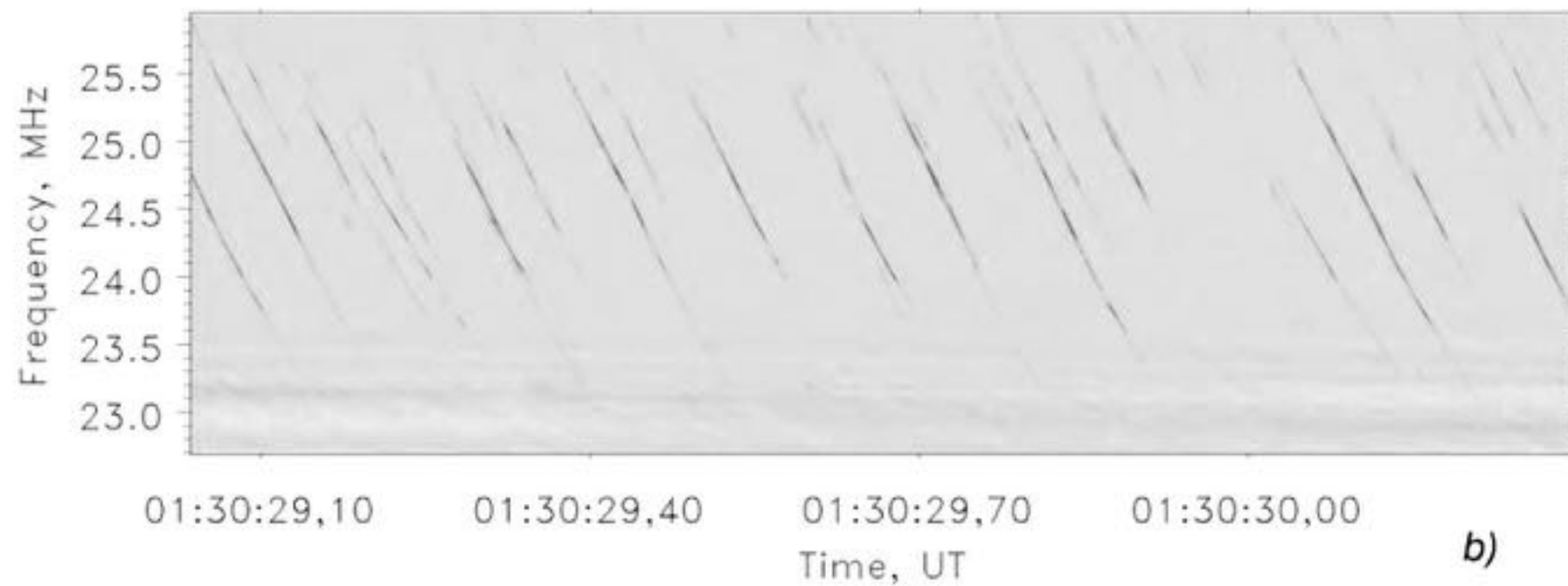
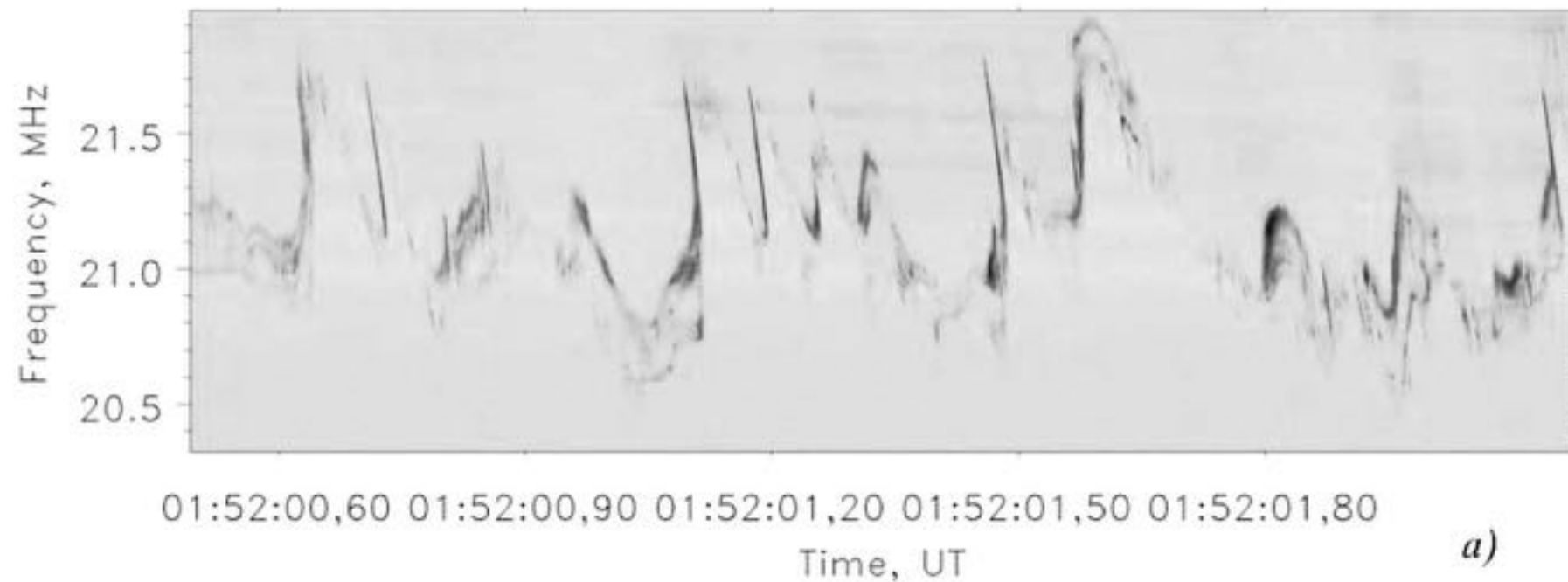
- digital SFA
- spectropolarimeter (4 Stokes)
- AOS

Jupiter: 1990-2000

- DAM polarization → N_e in sources [Dulk et al., 1992, 1994]
- Arcs morphology ? [Queinnec & Zarka, 1998]
- Correlations Radio/UV [Prangé et al., 1993]
- S-bursts fast measurements & scenario [Zarka et al., 1996]
- DAM microstructure [Carr et al., 1999]



Jupiter: 2000-2010



- digital (FFT) spectrographs at NDA & UTR-2
- waveform capture

Jupiter: 2000-2010

- Potential drops, e- & ion holes along Jupiter's B lines [Hess et al., 2007a, 2009]
→ e- acceleration by Alfvén waves, S-bursts & arcs modelling [Hess et al., 2007b, 2008a]
- DAM microstructure [Ryabov et al., 2007]
- First version of ExPRES code [Hess et al., 2008b]
- First correlations NDA-LOFAR [Nigl et al., 2007]

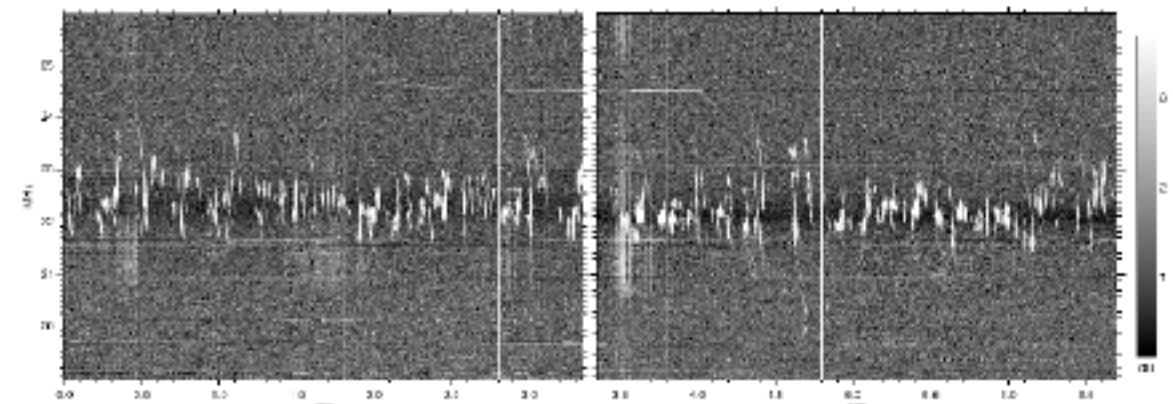
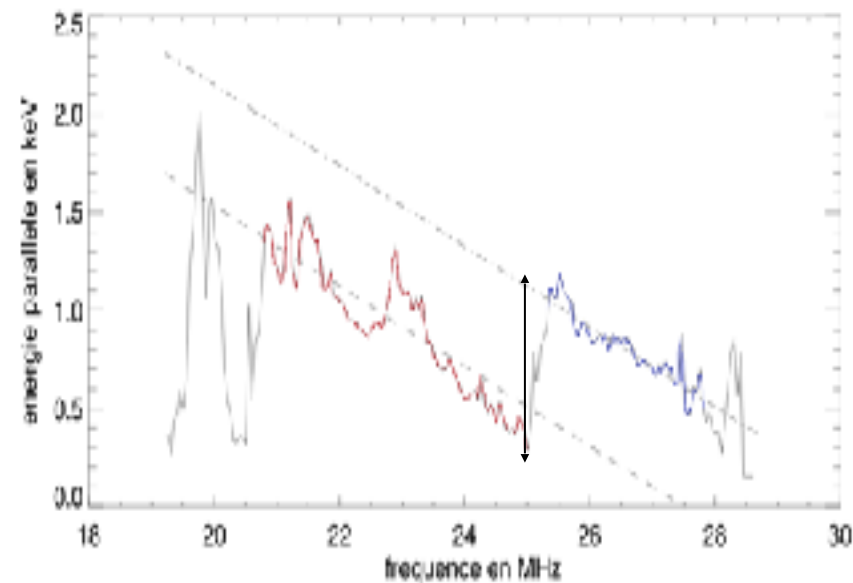
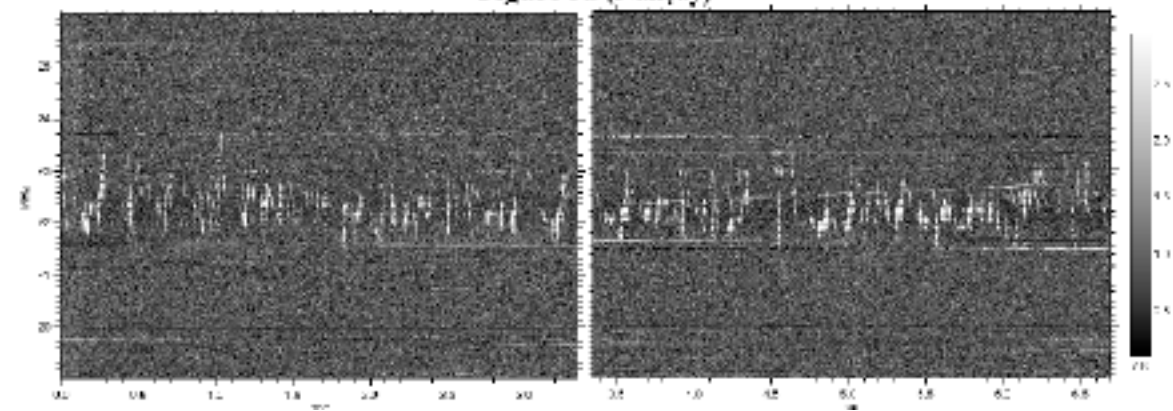
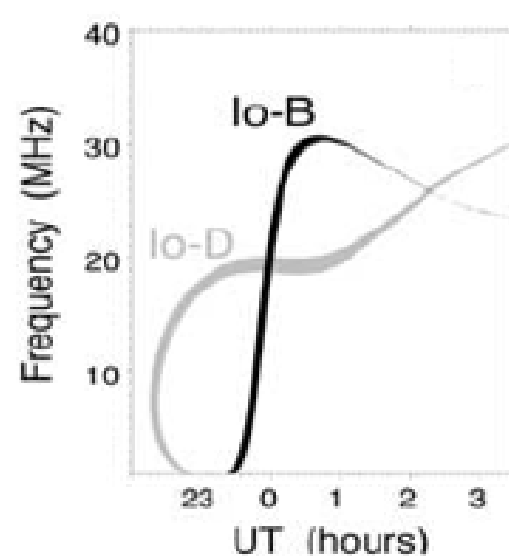
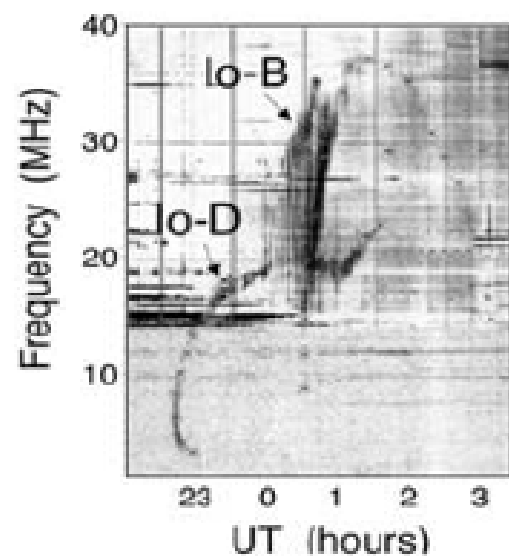
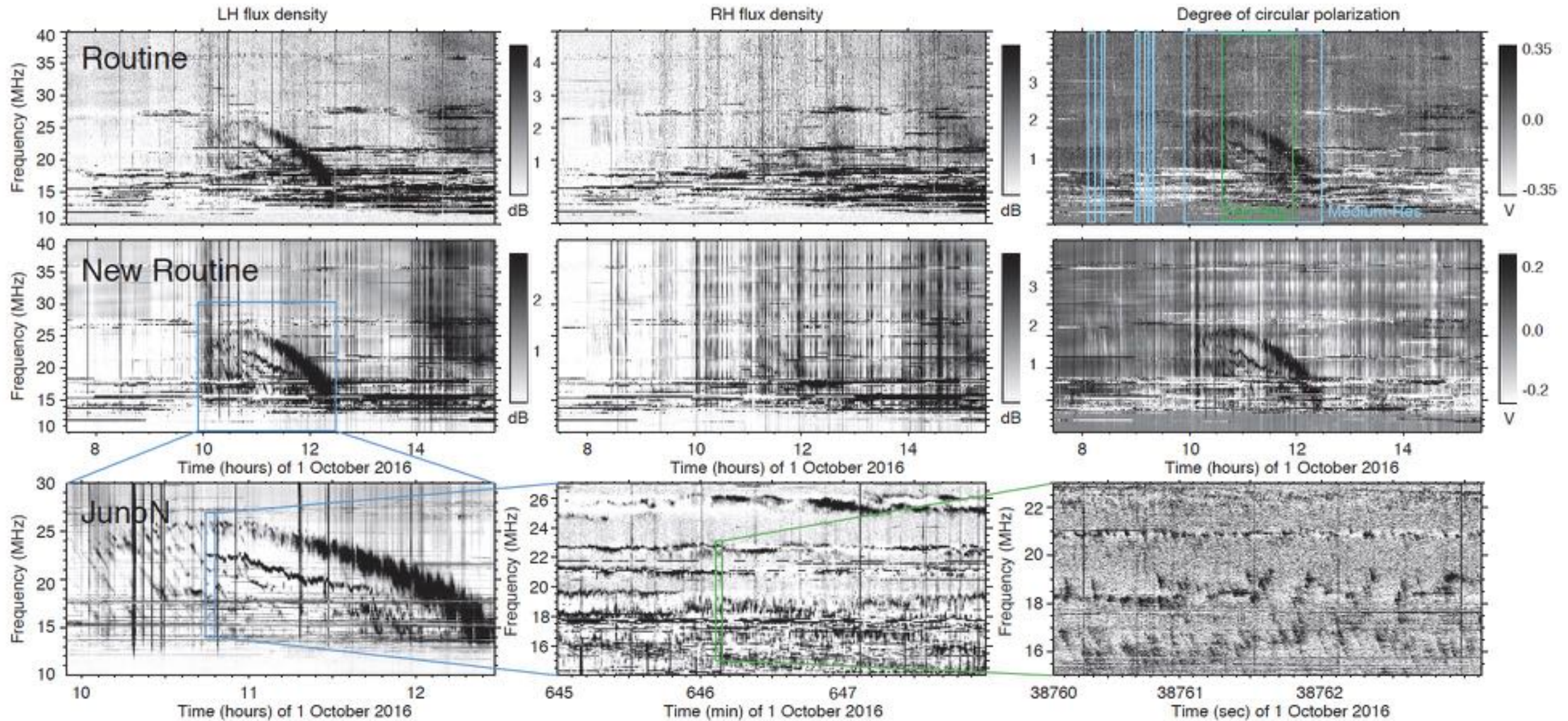


Figure 3a (Nançay)



Jupiter: 2010-2020

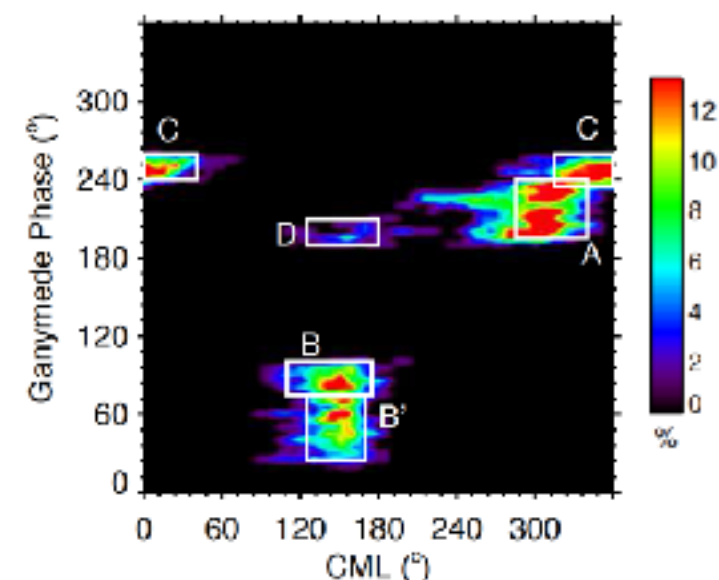
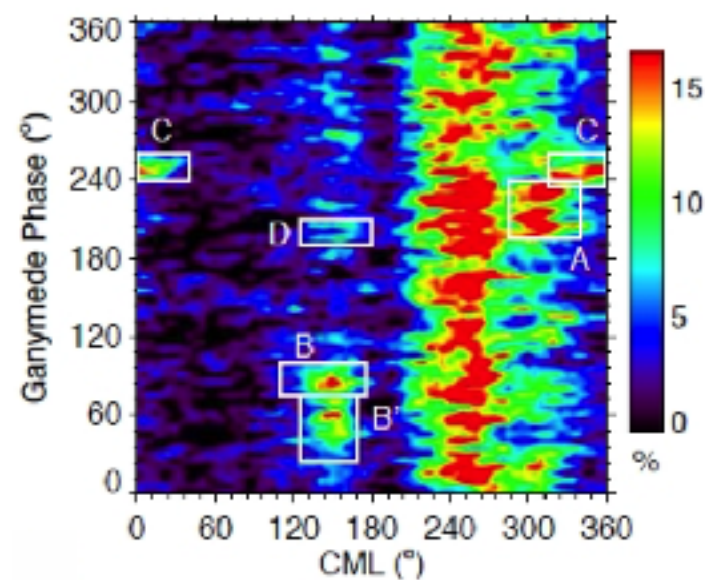
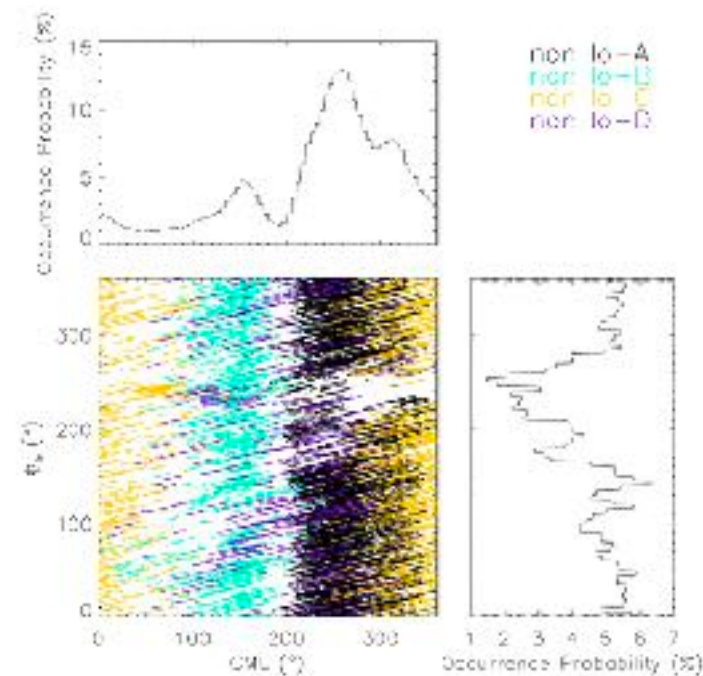
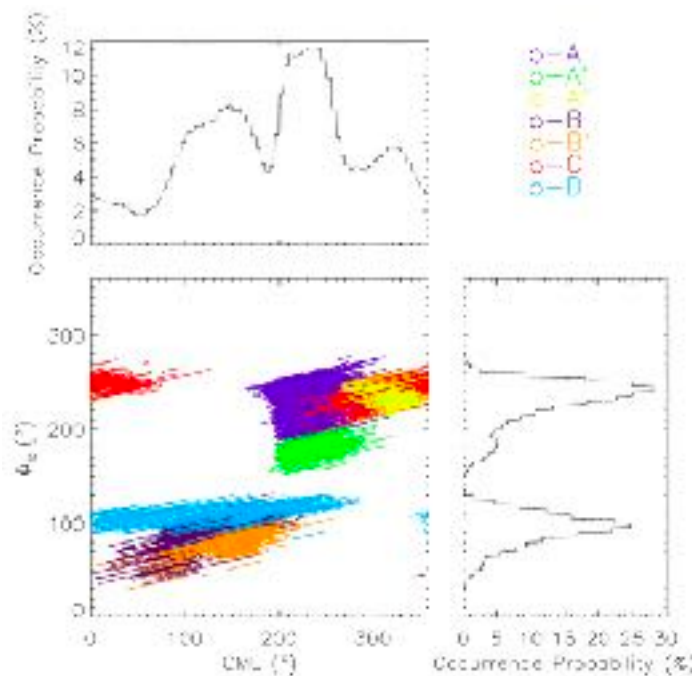


[Lamy et al., 2017]

→ NewRoutine, Mefisto & Juno-N receivers

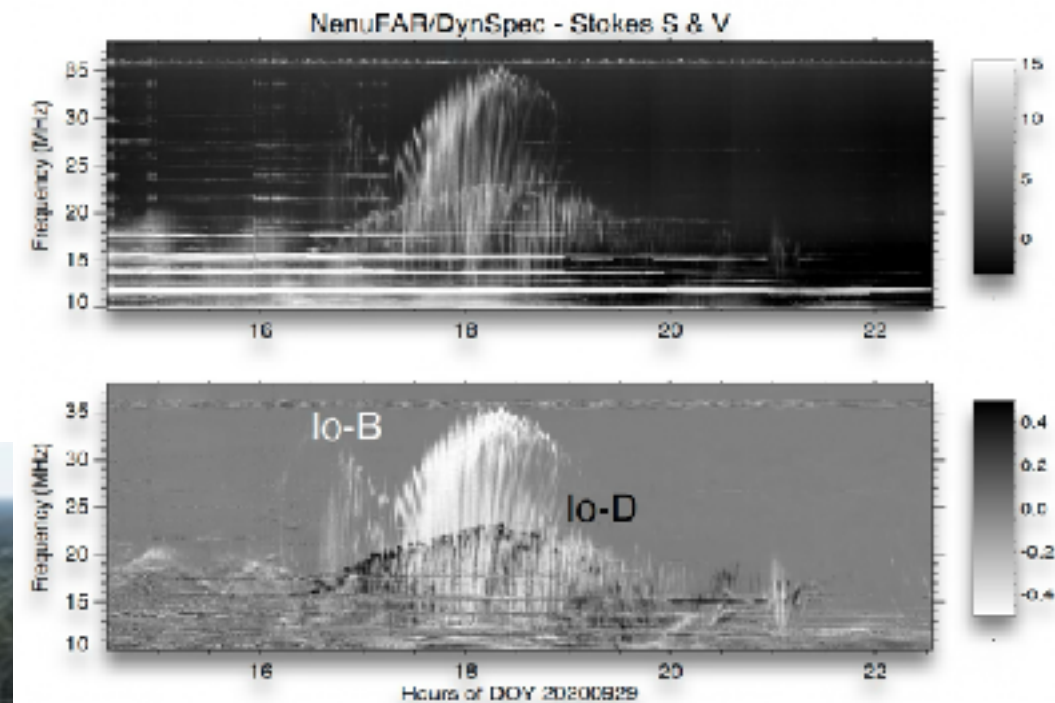
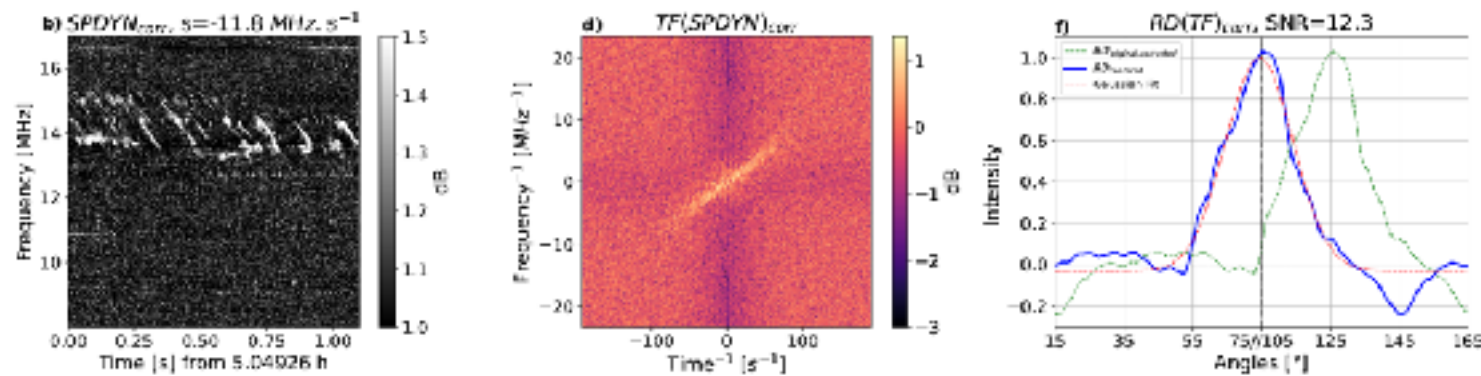
Jupiter: 2010-2020

- NDA catalog 1990-2016(2020) [Marques et al., 2017]
→ DAM control by Ganymede [Zarka et al., 2018]
- Final version of ExPRES code [Louis et al., 2019]
→ First predictions for Exoplanets [Hess & Zarka, 2011]



Jupiter: 2020+

- Digitization of NDA analog archives [Lamy et al., 2023]
- DAM control by Europa [Jacome et al., 2022]
- Automatic recognition of S-bursts [Mauduit et al., 2023]
- NenuFAR beamformed studies of Jupiter emissions [Lamy et al., 2022]



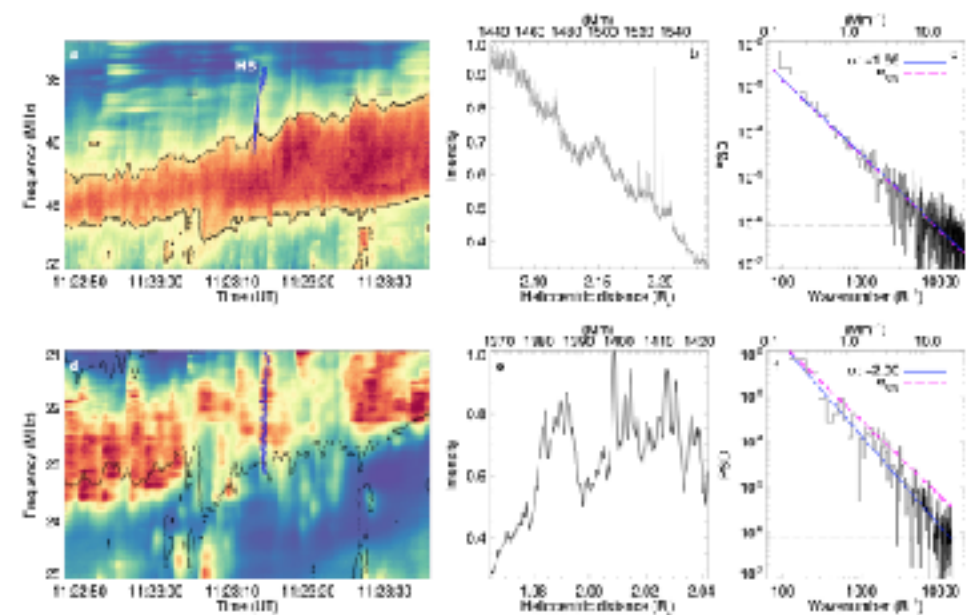
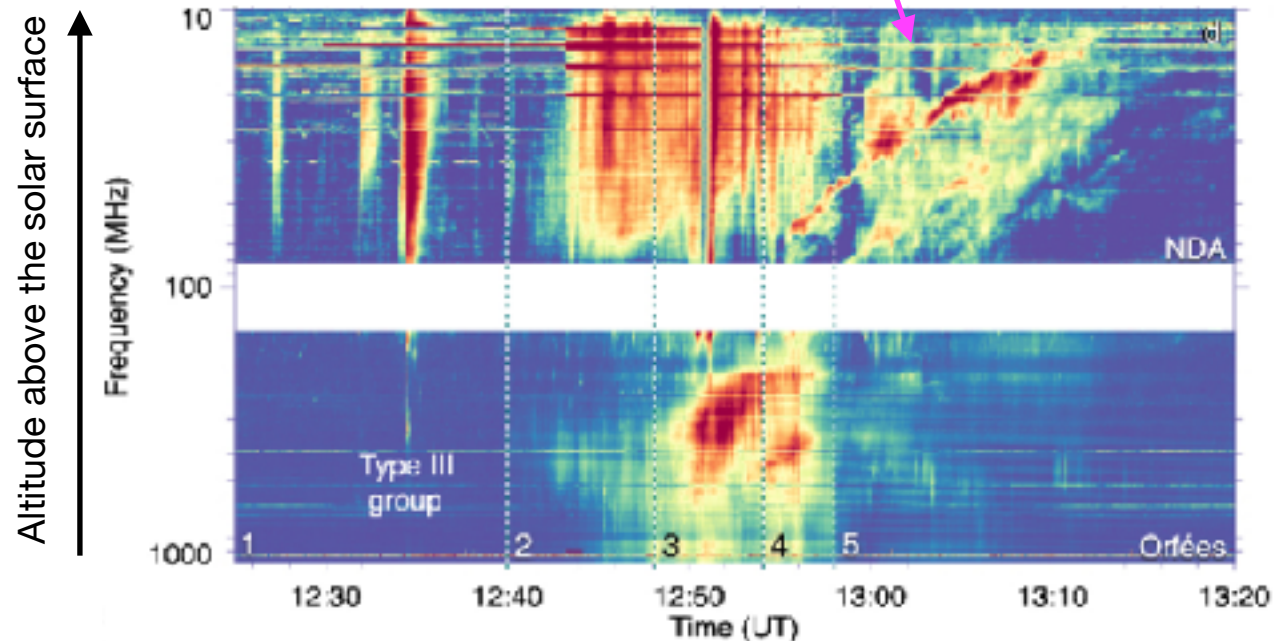
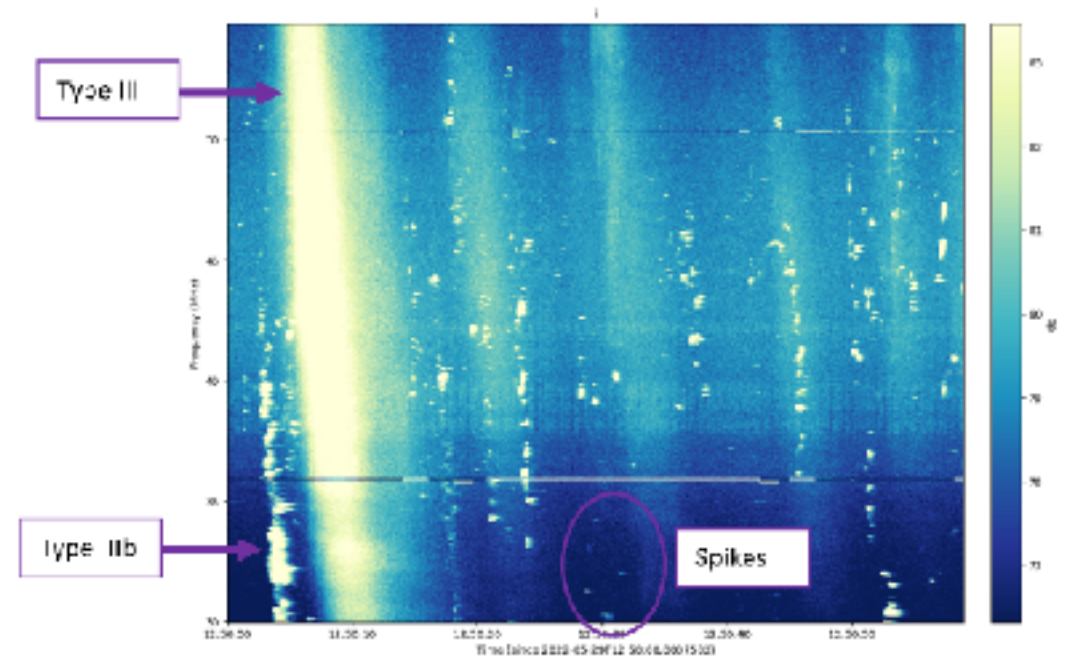
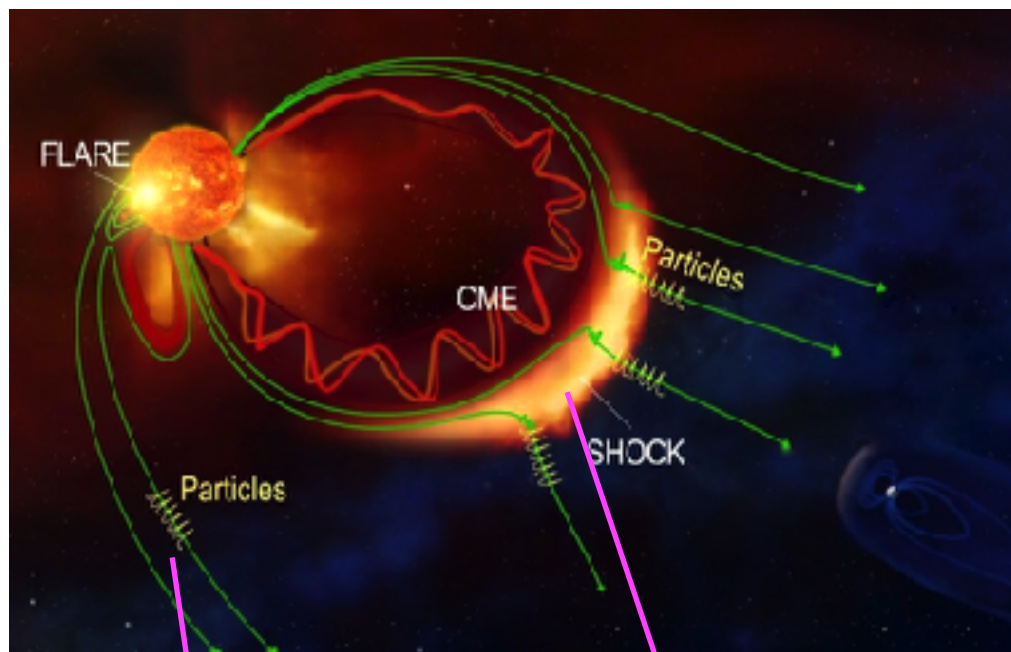
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Sun

- Solar burst studies (type II & III & spikes)
- NenuFAR beamformed studies of Solar bursts

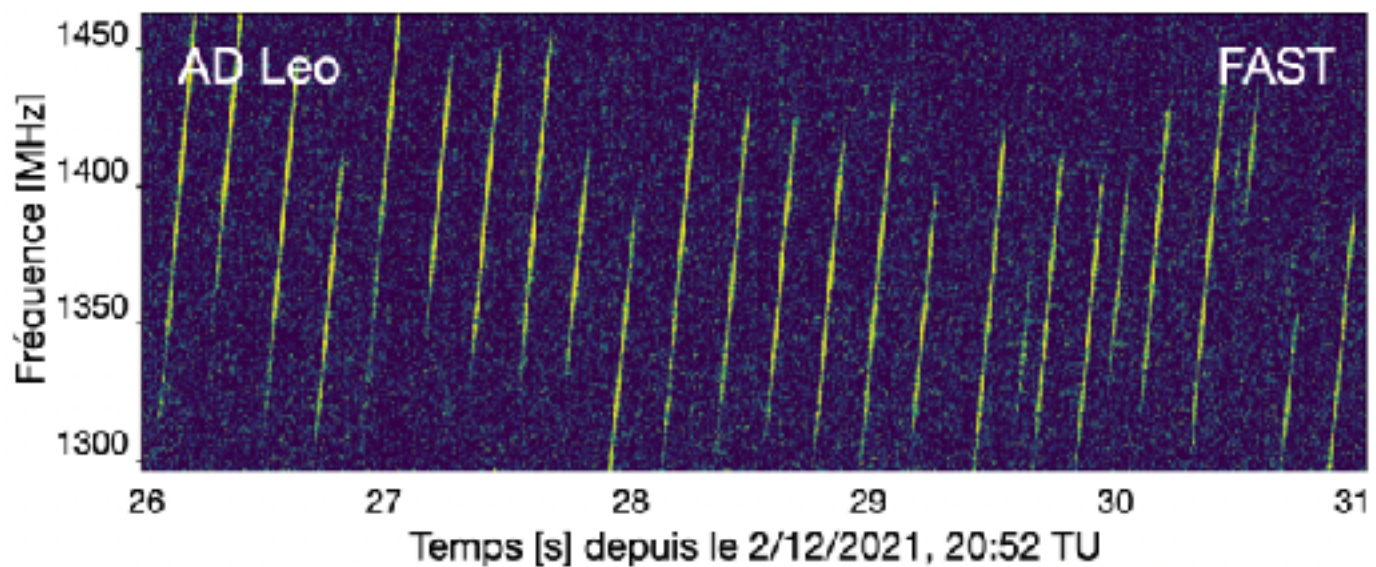
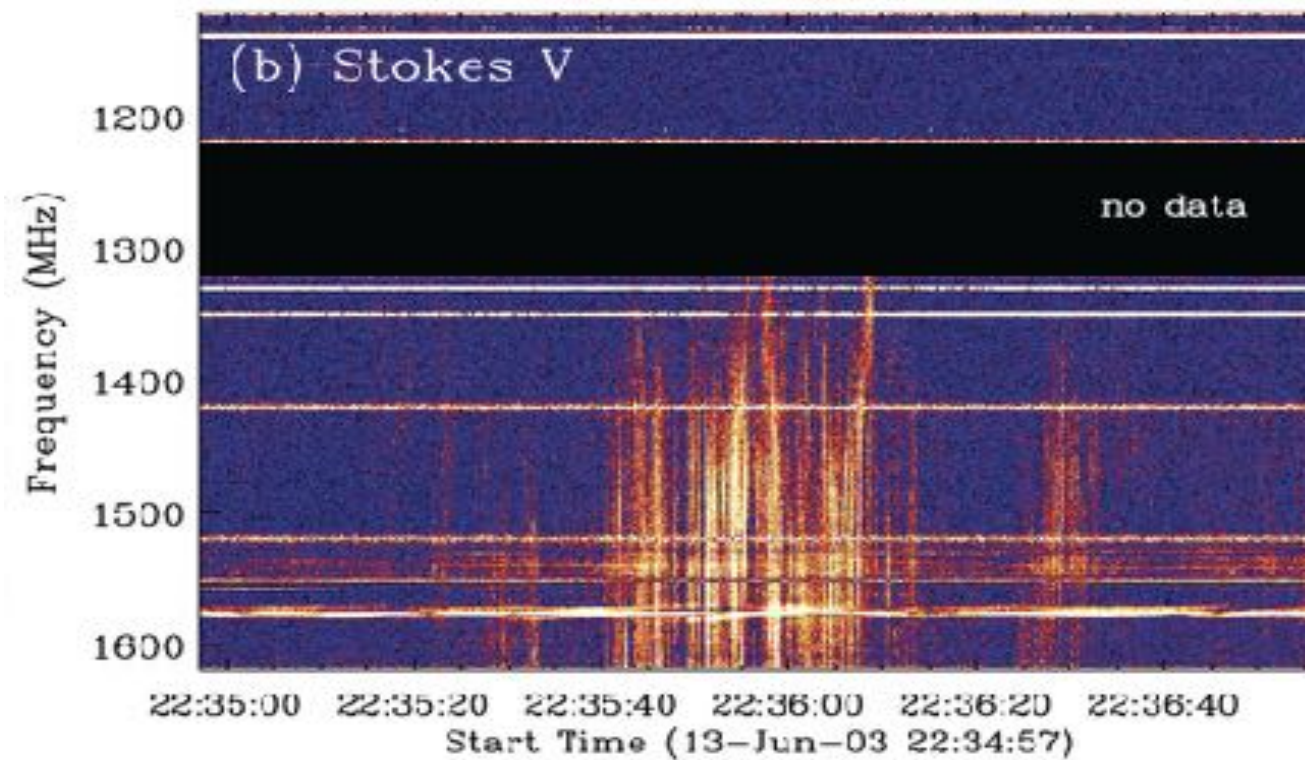
[Briand et al., 2008; Melnik et al., 2015; Pick et al., 2016; Shevchuk et al., 2016]

[Carley et al., 2021; Briand et al., 2022]



Stars

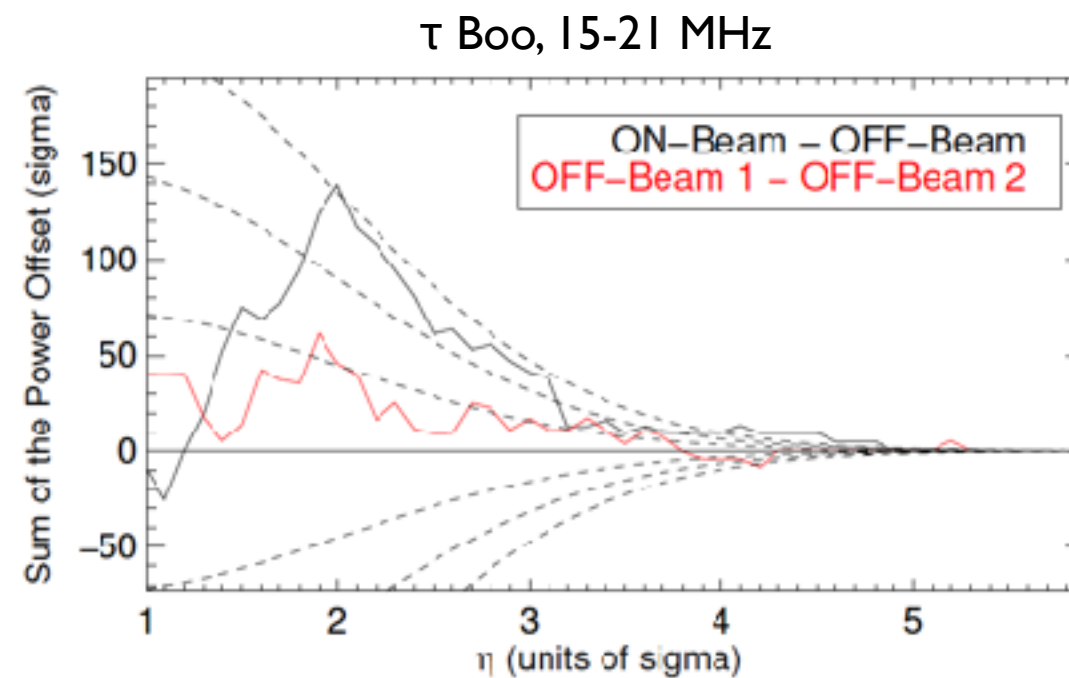
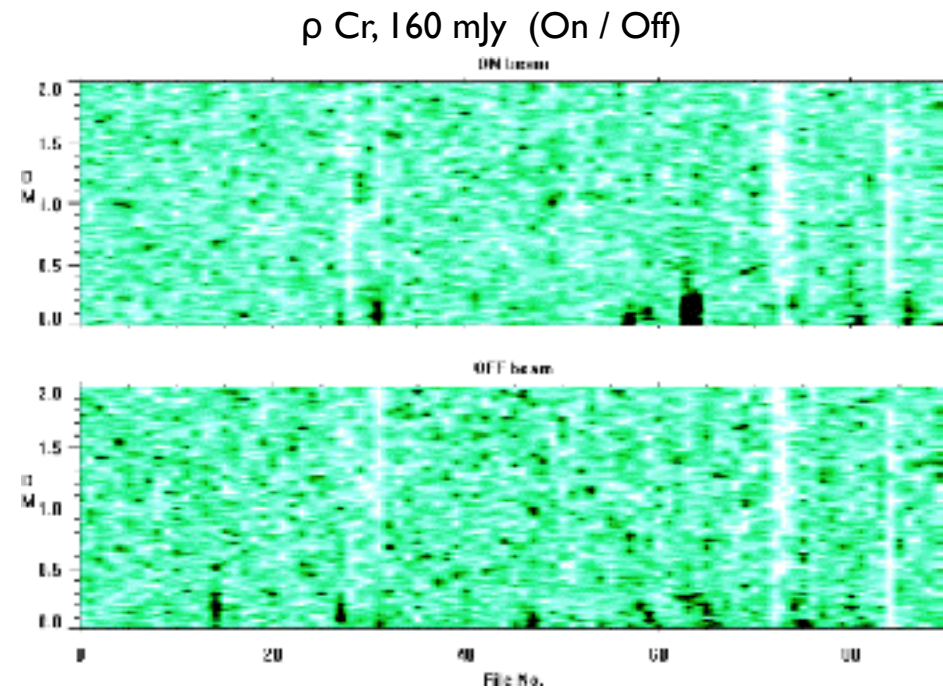
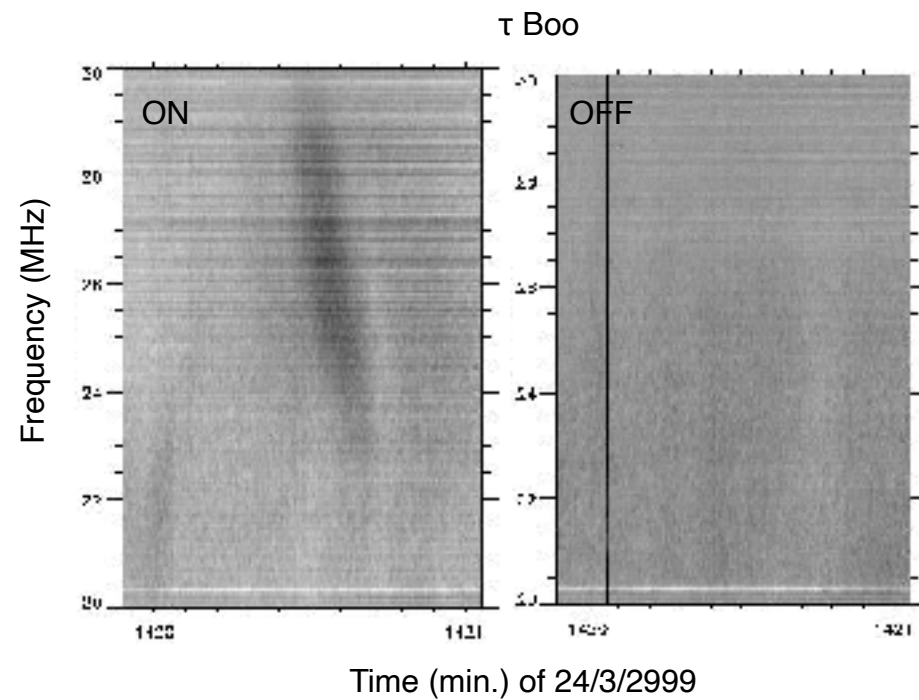
- Bursts, fine structures, light curves [Osten et al., 2006, 2008; Zhang et al., 2023; Zarka et al., 2025]



Exoplanets

- Inconclusive tentative detections

[Ryabov et al., 2004; Vasylieva, 2015; Turner et al., 2021, 2024]



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Interferometers: imaging studies



VLA

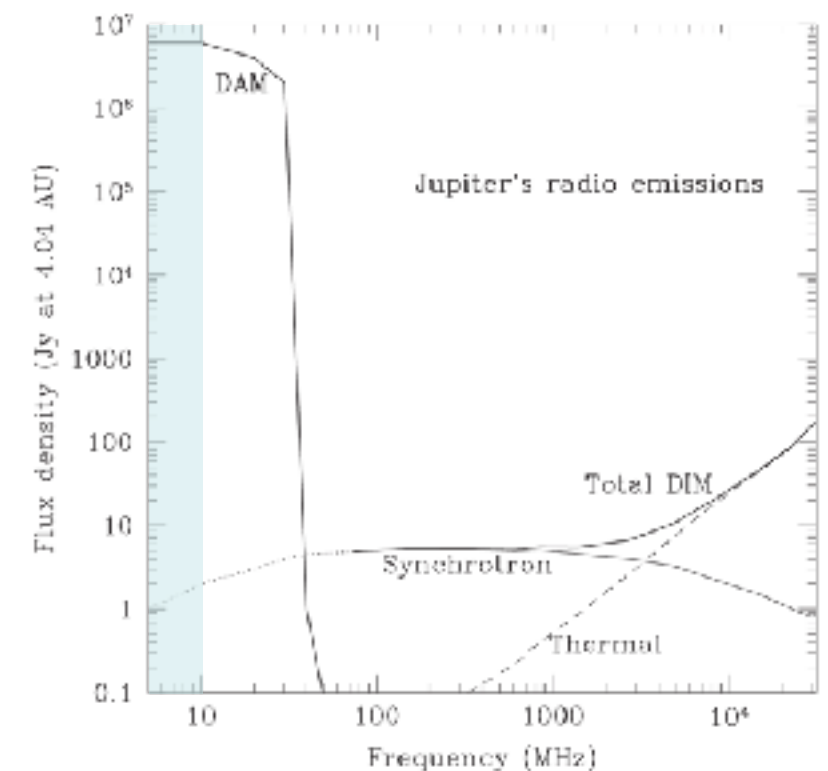
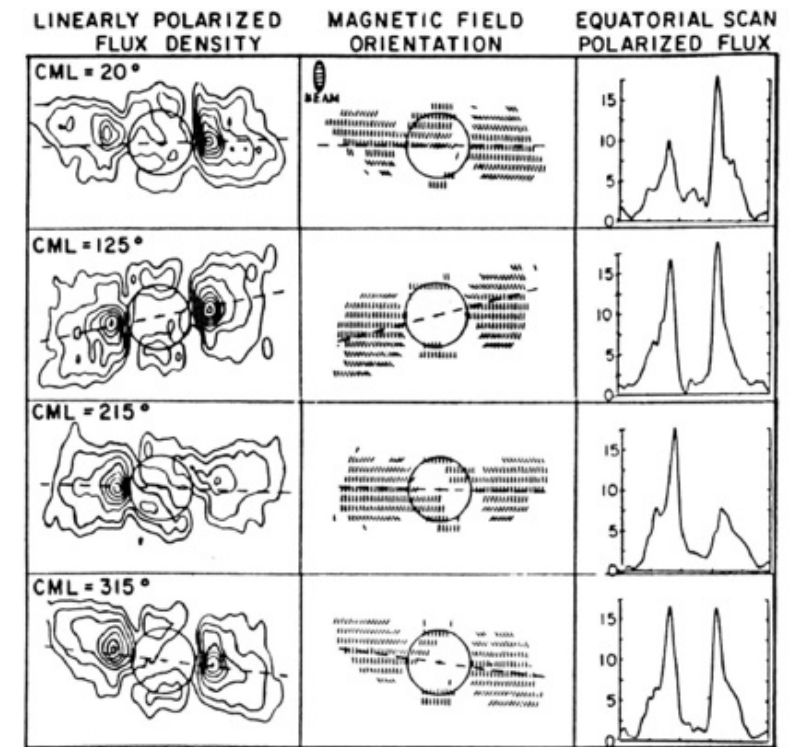
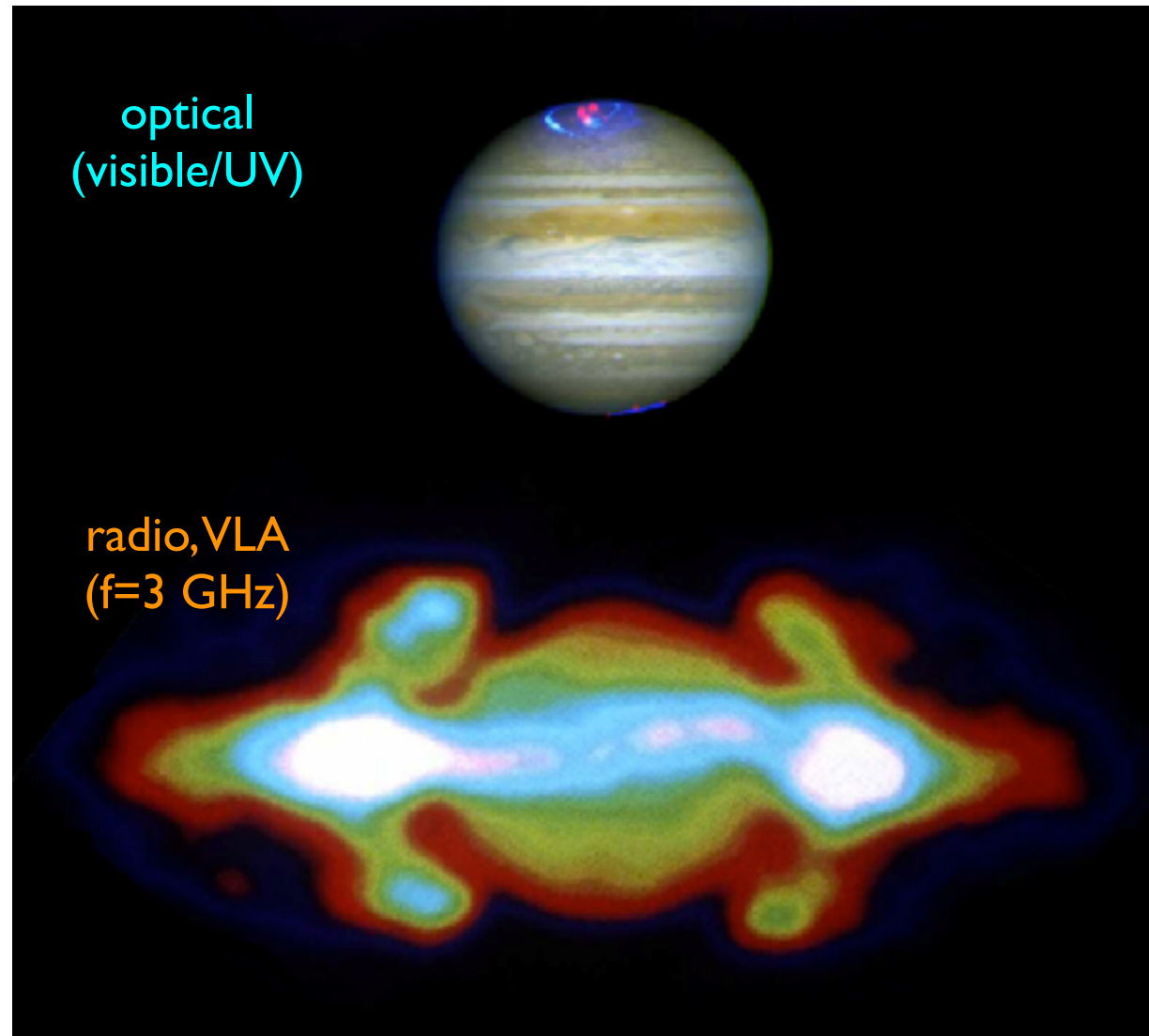


LOFAR

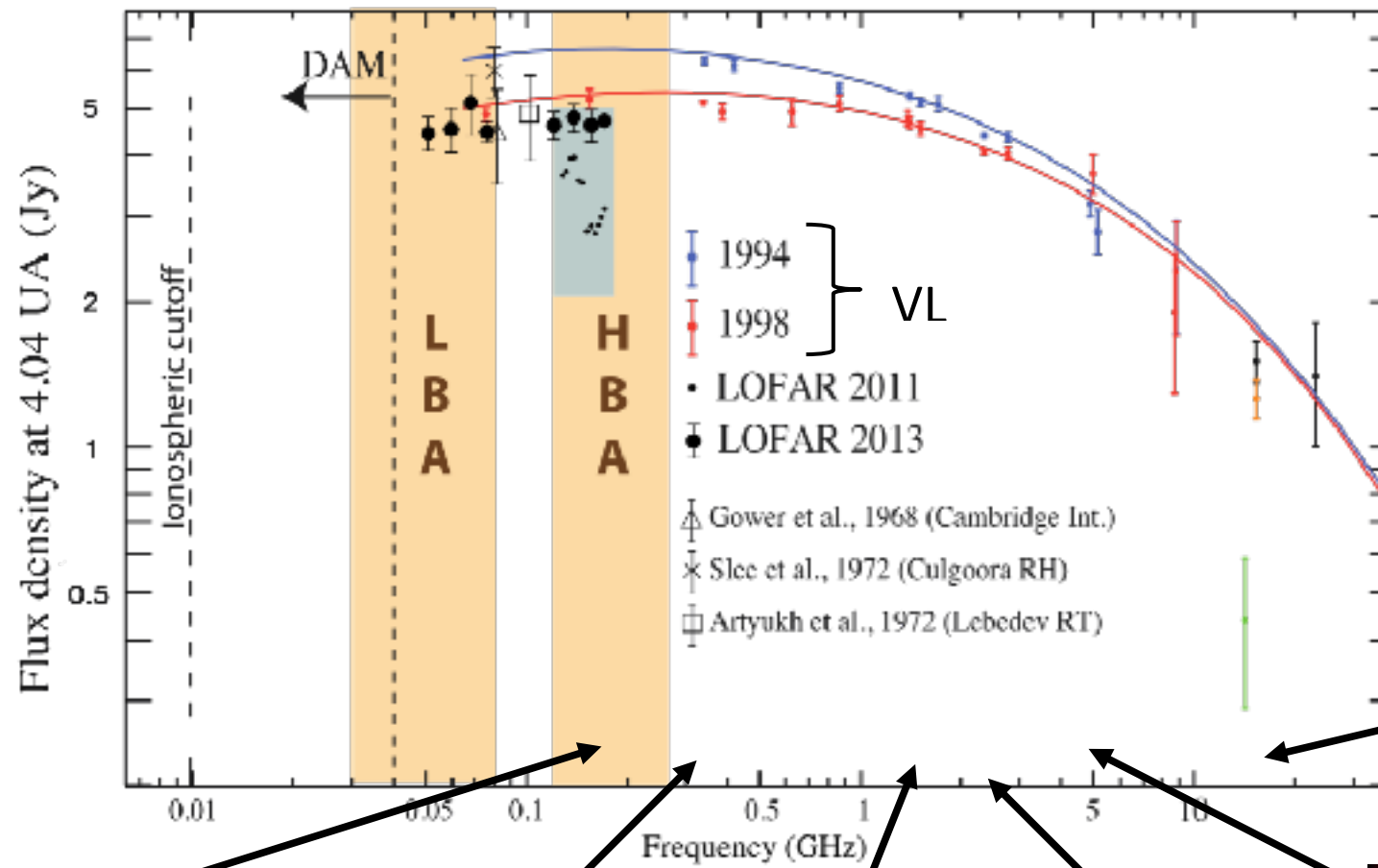
Jupiter: imaging

- Imaging, polarization, variability of radiation belts (incl. SL9 impact)

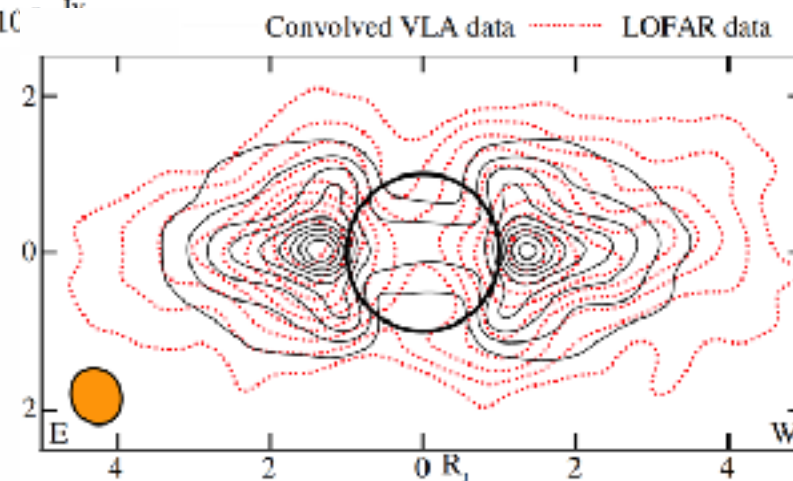
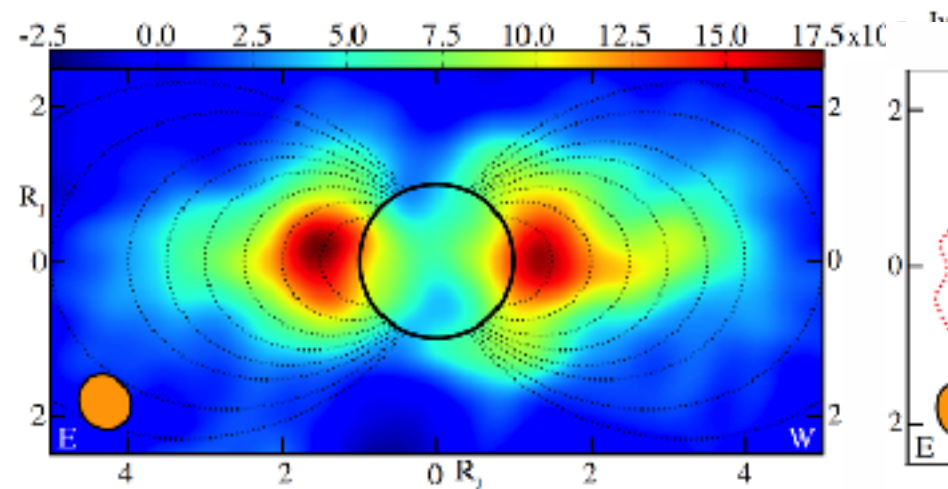
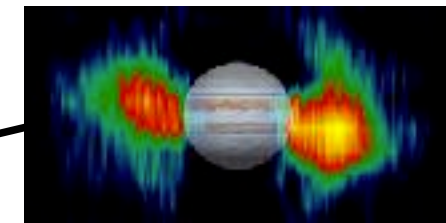
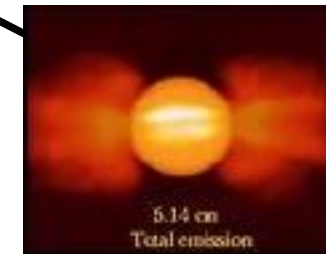
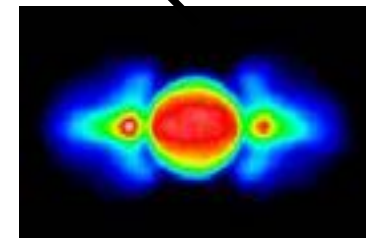
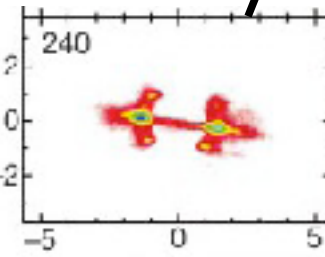
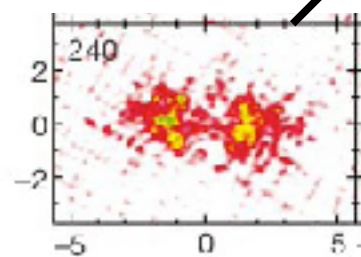
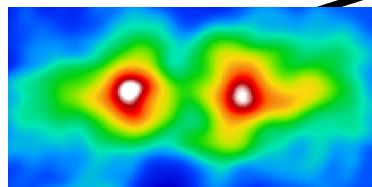
[de Pater et al., 1981...2004; Santos-Costa et al., 2001...2014]



Jupiter: imaging

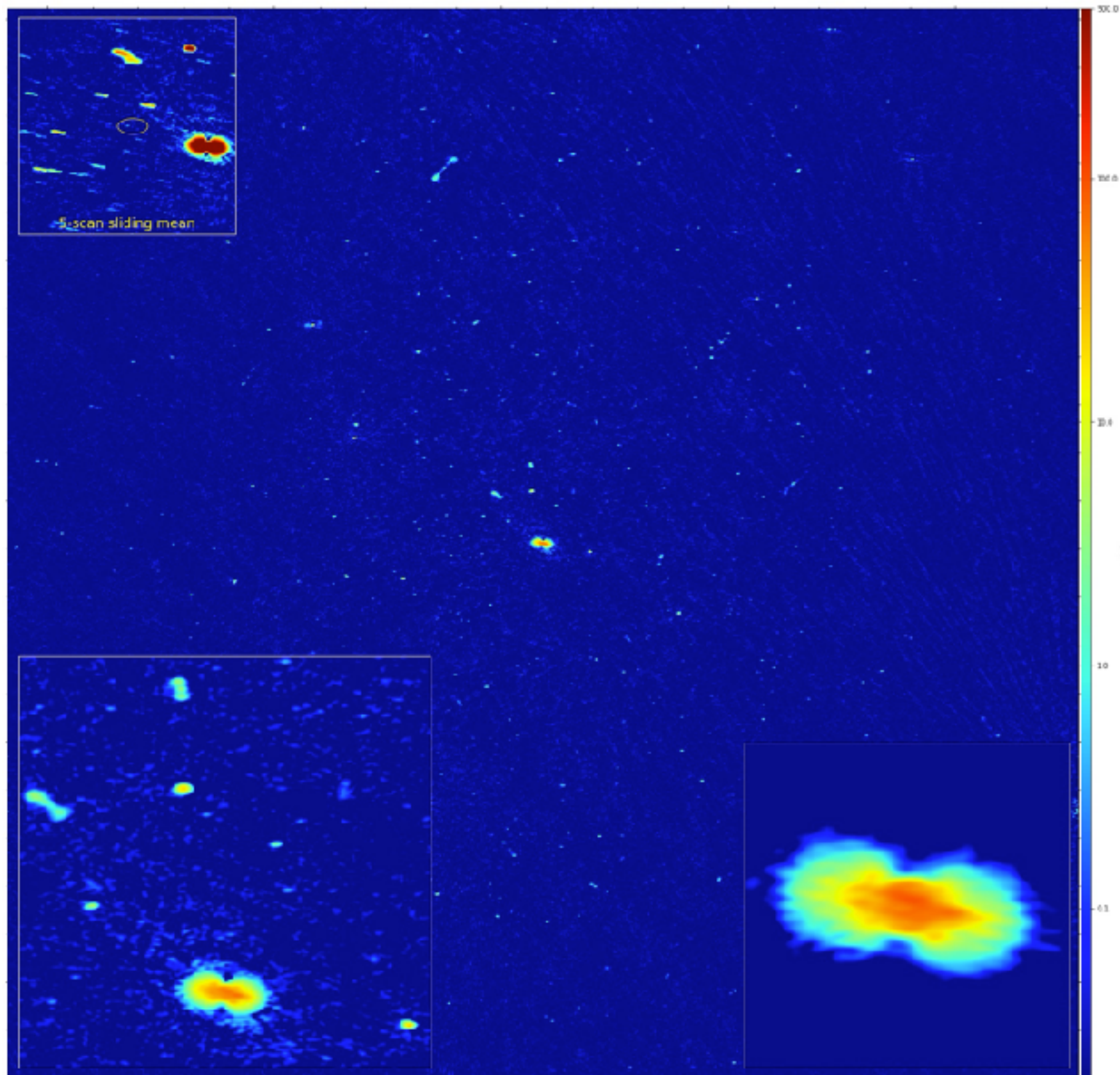


[de Pater & Dunn, 2003]



[Girard et al., 2016]

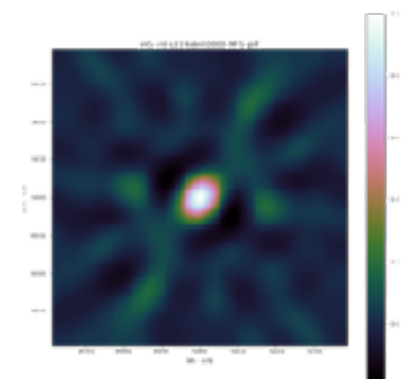
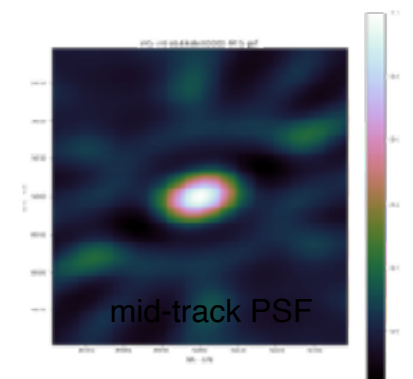
Jupiter: imaging



MeerKAT movie

Great Conjunction of
21/12/2020
880 - 1650 MHz
~17 min. / image + gaps
27 images : total ~8h30

Credits : Oleg Smirnov &
RATT (ratt.center) &
SARAO (sarao.ac.za)



end-track PSF

Jupiter : VLBI

- Limits on DAM source size [Dulk, 1967]
- VLBI on DAM [Wucknitz et al., 2024]

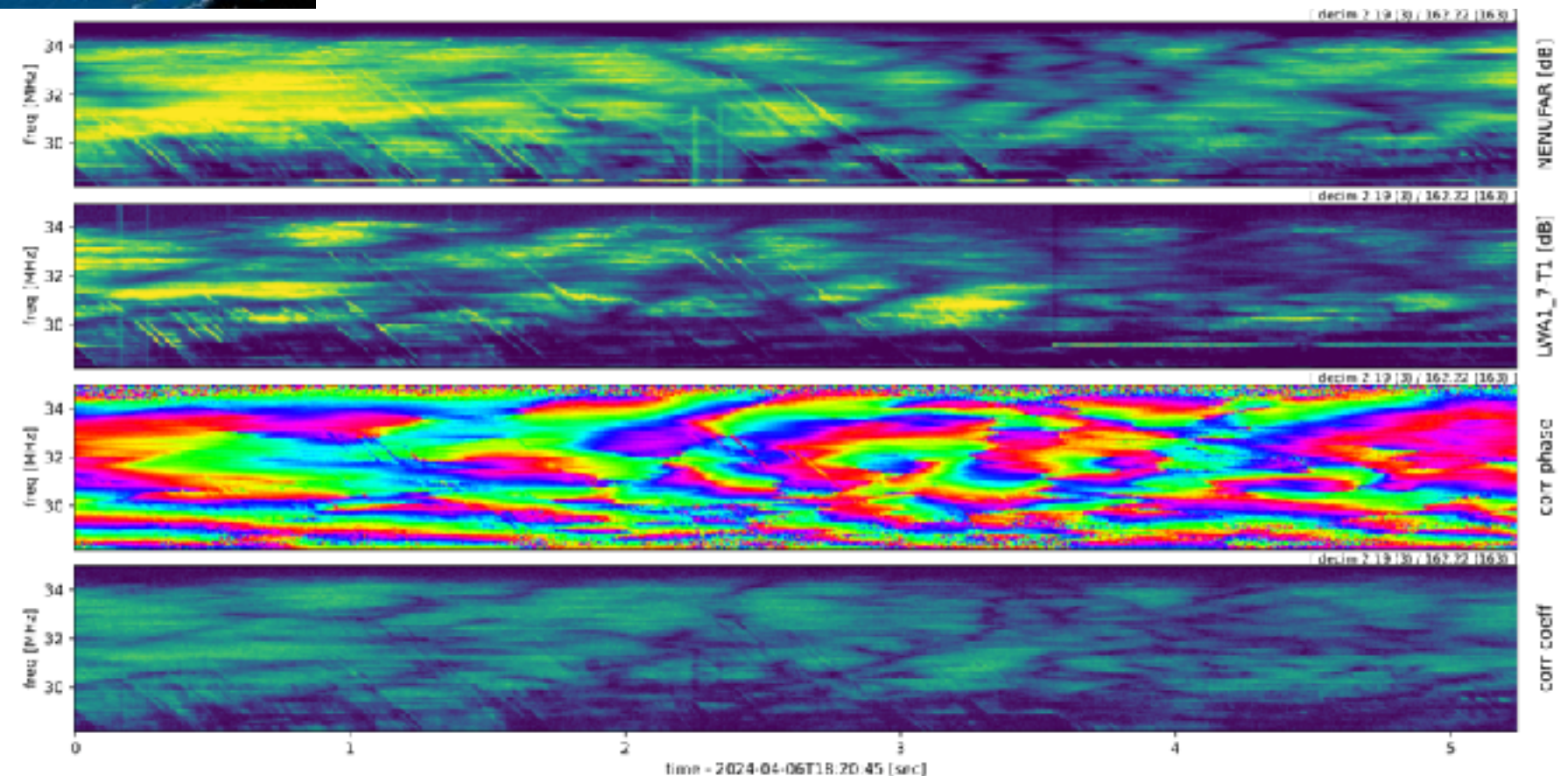
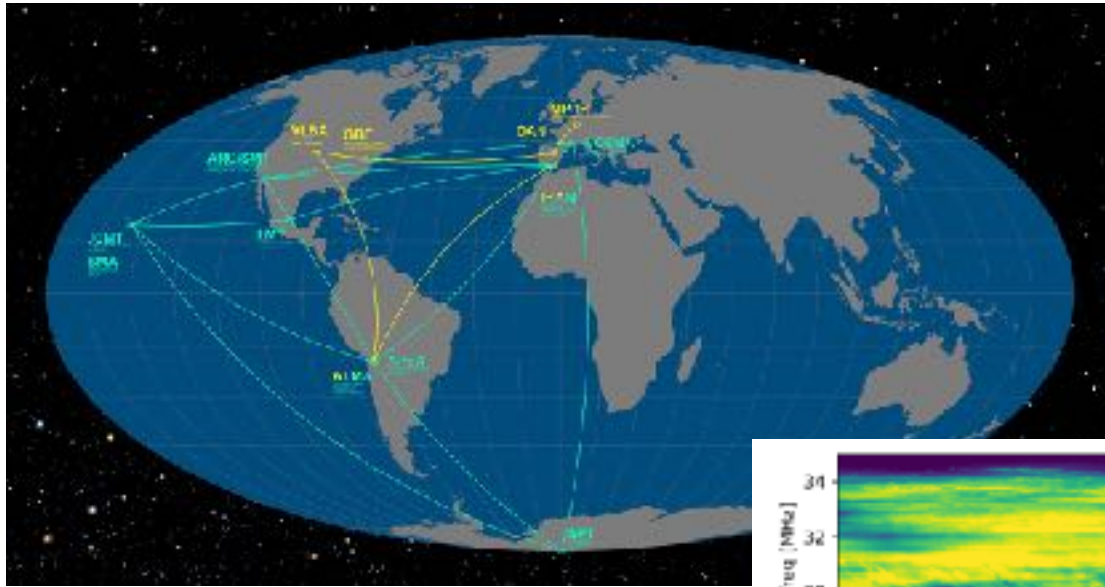


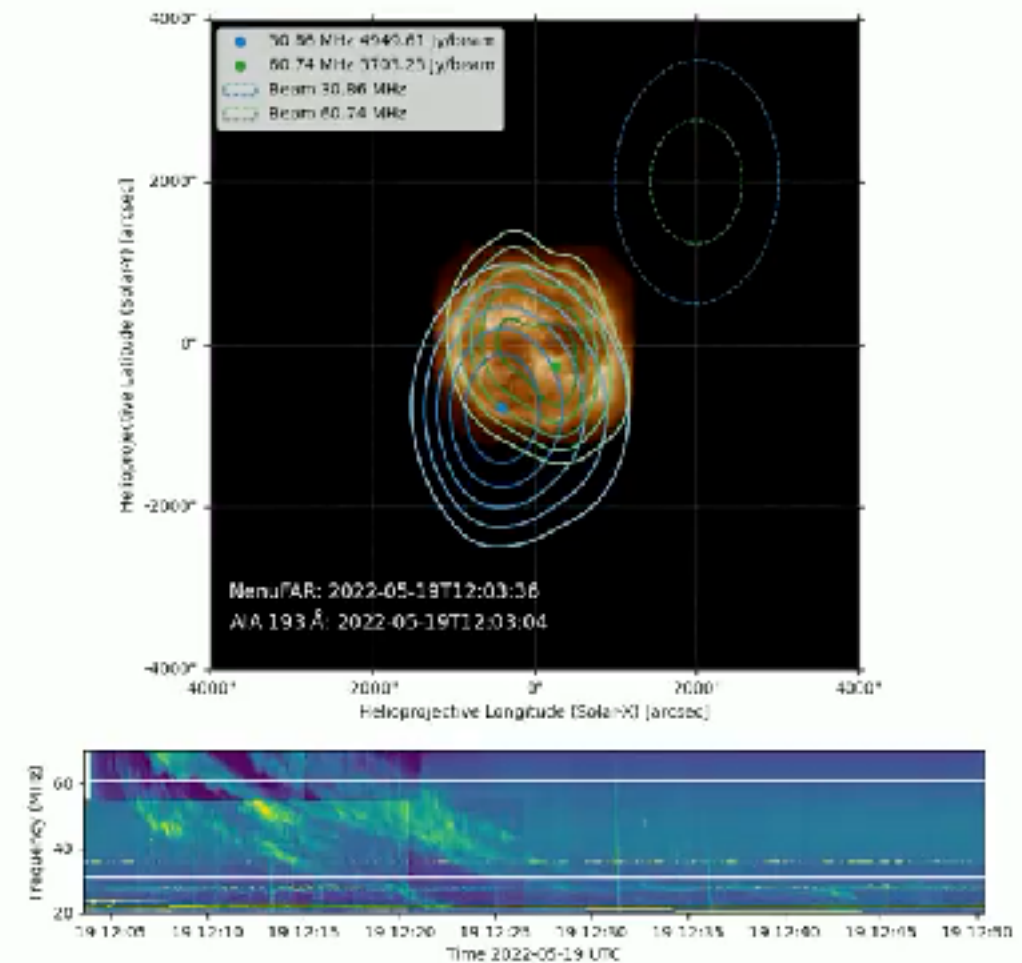
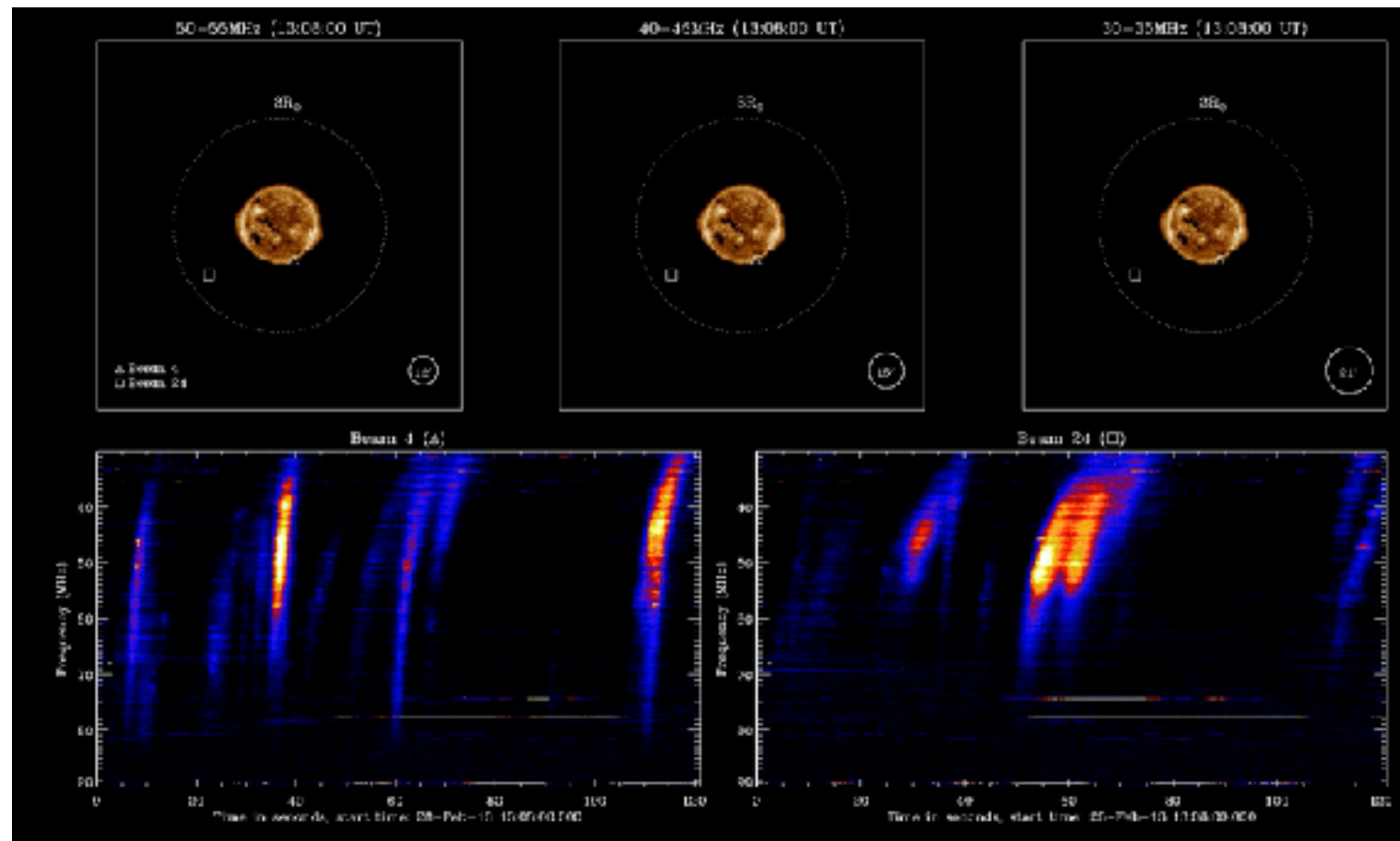
Fig. 7. Fringes on the transatlantic NenuFAR–LWA1 baseline, panels as in Fig. 5. To our knowledge, these are the first transatlantic fringes at 30 MHz.
(Wucknitz et al., 2024)

Sun: imaging

- Solar multi-beam studies with LOFAR
- NenuFAR imaging & t-f analyses

[Morosan et al., 2014, 2015, 2017, 2019, 2022, 2025]

[Briand, Murphy...]



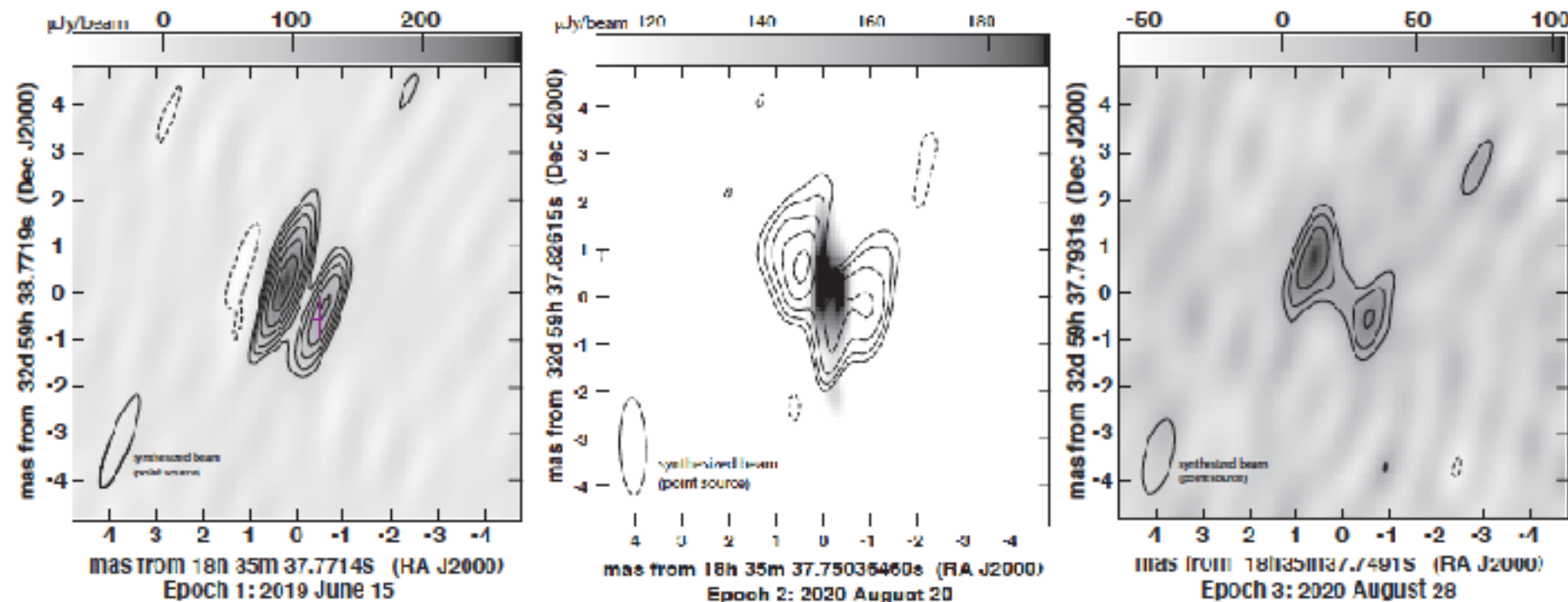
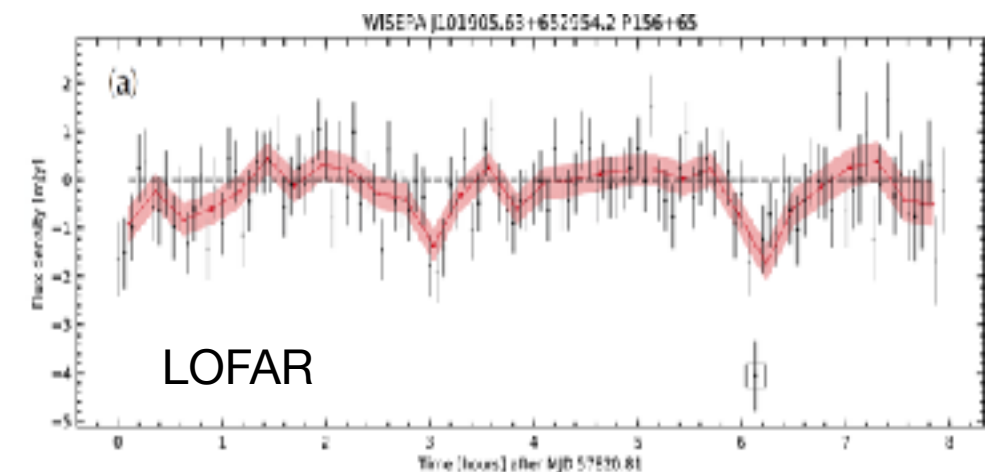
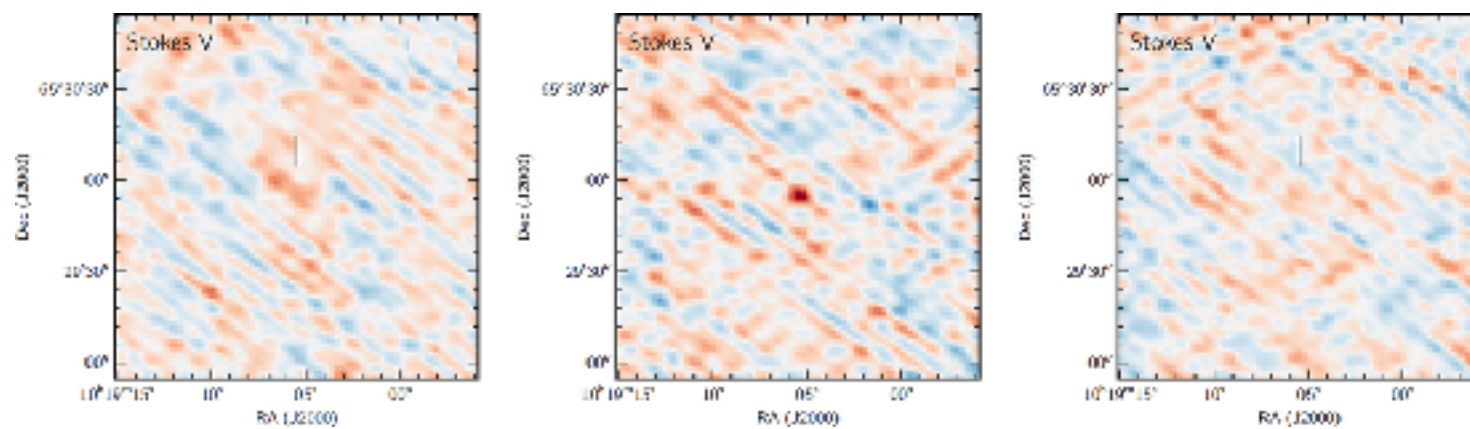
Stars: imaging & VLBI

- Stellar observations & discoveries with LOFAR/LoTSS → BD, T Tau, RSCVn...

[Coughlan et al., 2017, Vedantham et al. 2020,2022,2023; Callingham et al, 2021a,b; Davis et al., 2021; Tet et al., 2021; Feeney-Johansson et al., 2021; Hajduk et al., 2022]

- Stellar radiation belts (HSA, EVN) → mas, 0.1 mJy

[Kao et al., 2023 ; Climent et al., 2023]

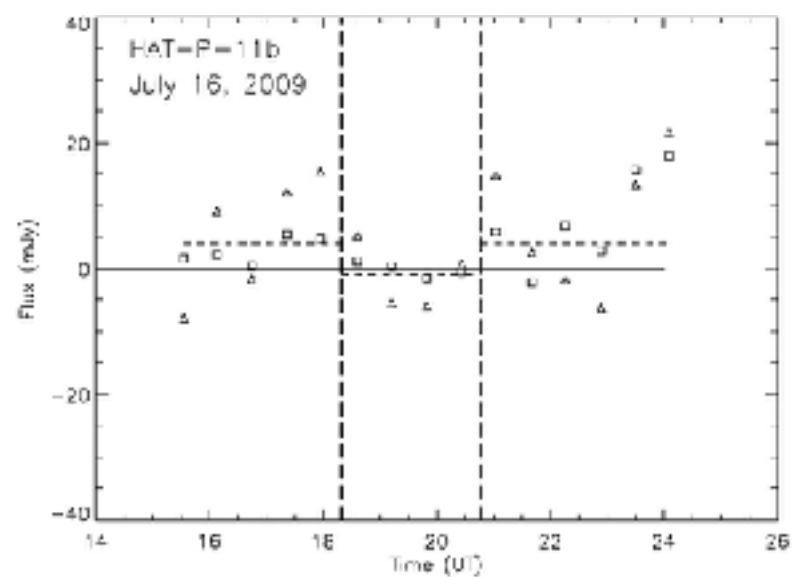


Exoplanets: imaging

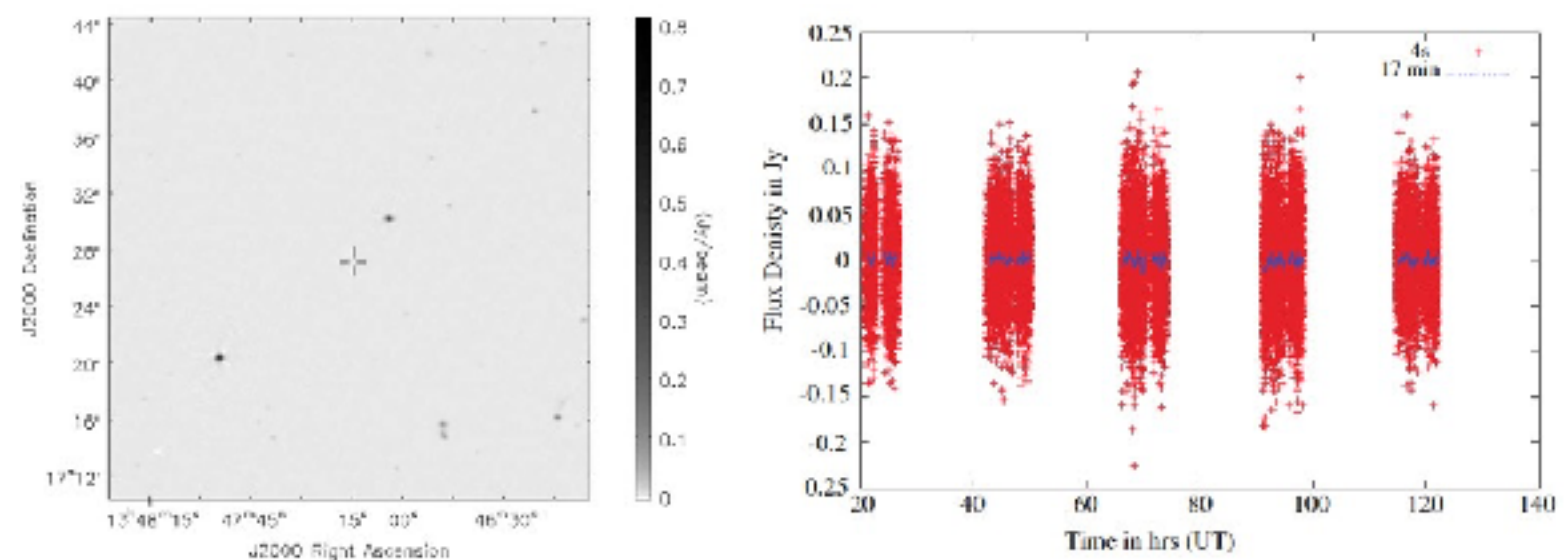
- Clues, candidates, but no confirmed detections (GMRT, VLA, LOFAR)

[Lecavelier et al., 2013; Hallinan et al., 2013; deGasperin et al., 2020, Vedantham et al. 2020]

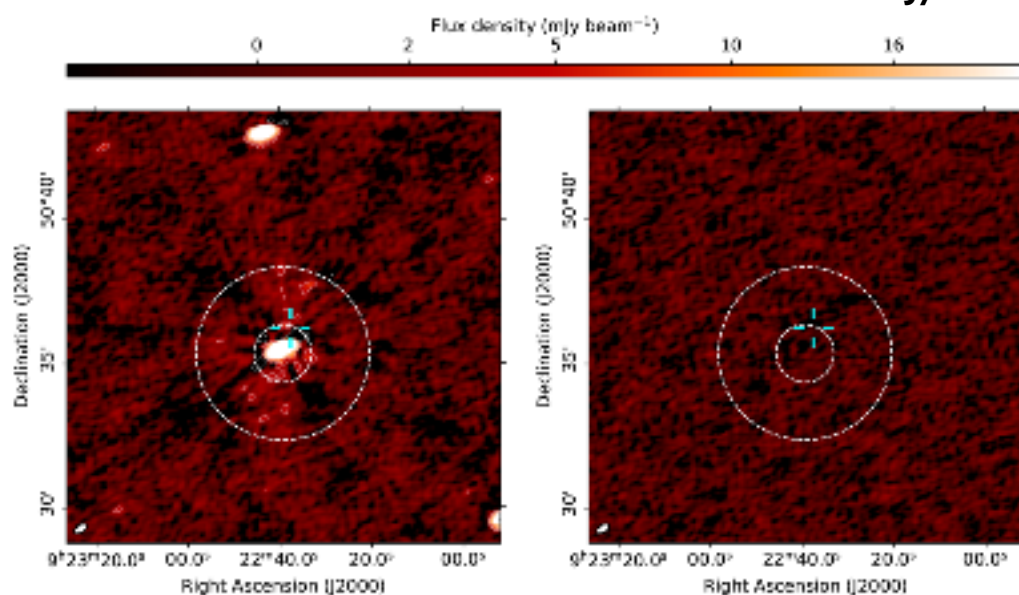
GMRT, Hat-P-11, 150 MHz, ~4 mJy



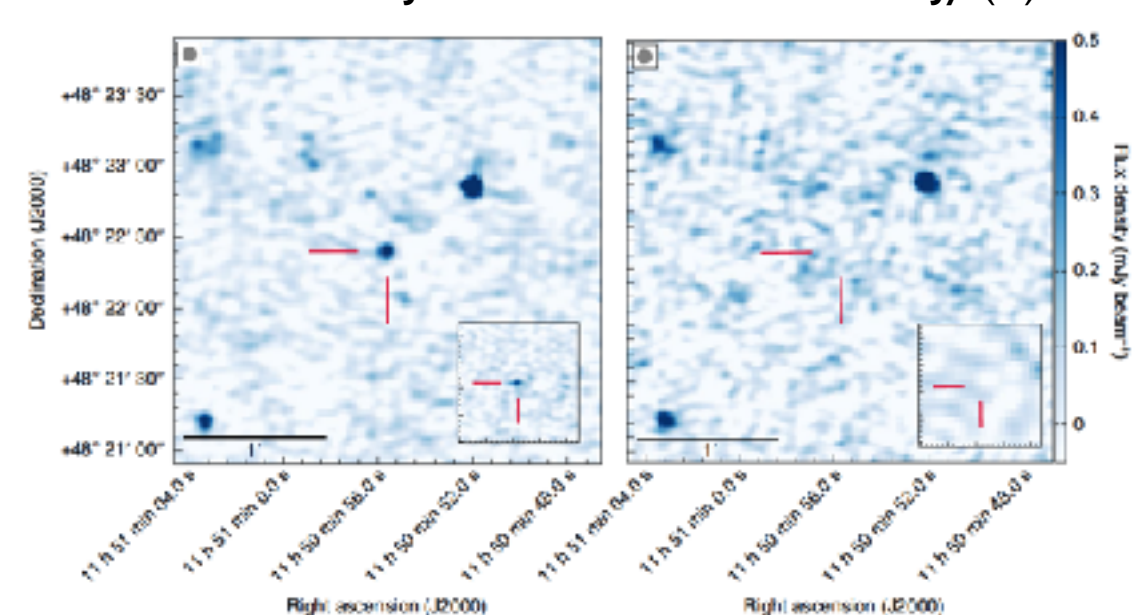
GMRT, τ Boo, 150 MHz, <1 mJy



LOFAR, HD 80606, 30-78 MHz, <10 mJy

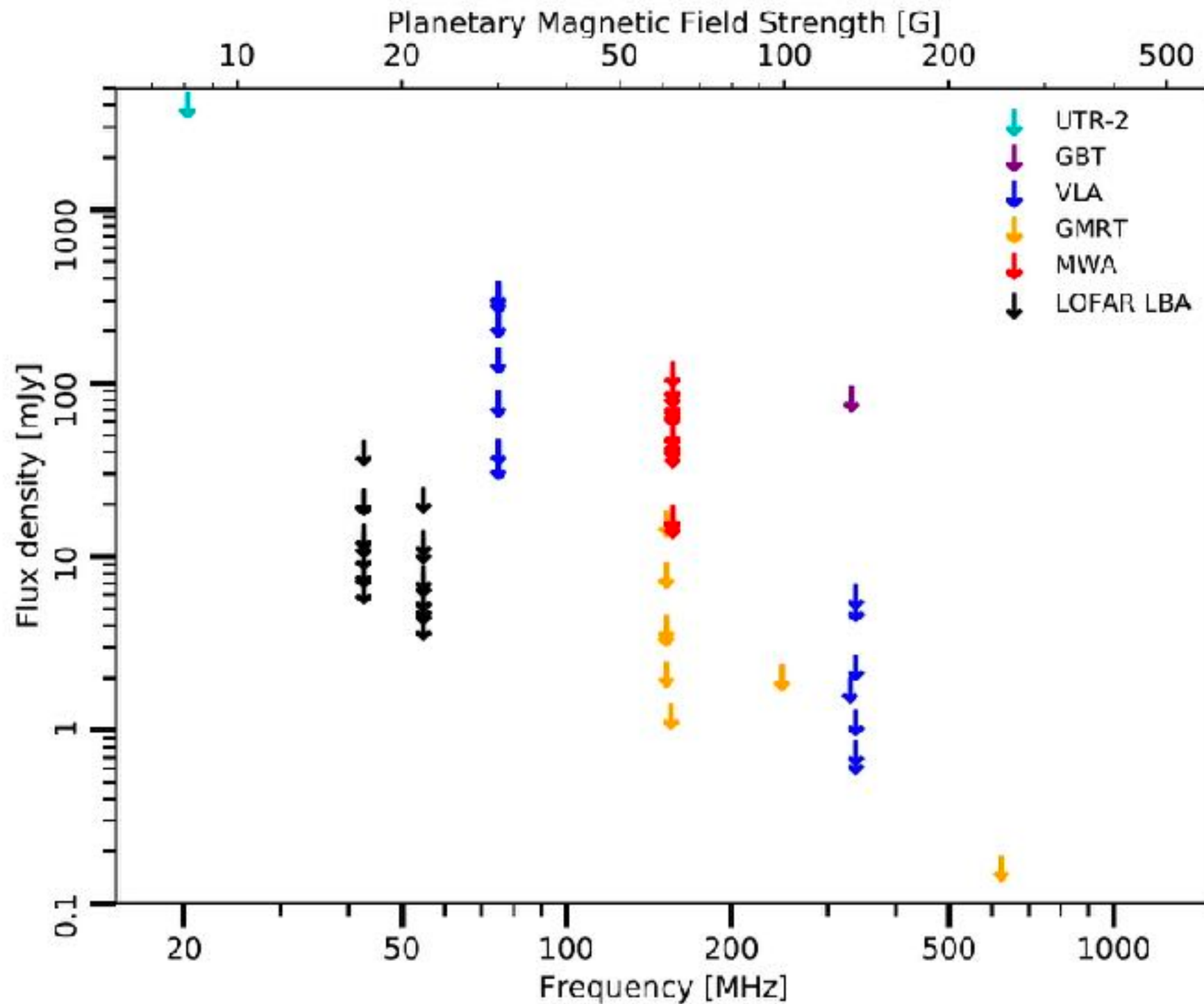


LOFAR, GJ 1151, 120-170 MHz, ~1 mJy (V)



Exoplanets: imaging

→ Where are they ?



- Heroic times
 - Jupiter, Sun
- Single dishes and phased arrays: t-f studies
 - Jupiter (<1990, 1990-2000, 2000-2010, 2010-2020, 2020+)
 - Sun, Stars, Exoplanets
- Interferometers: imaging studies
 - Jupiter, Sun, Stars, Exoplanets
- Dynamic spectra (t-f) from imaging data
 - Stars, Exoplanets
- Prospects

Dynamic spectra (t-f) from imaging data

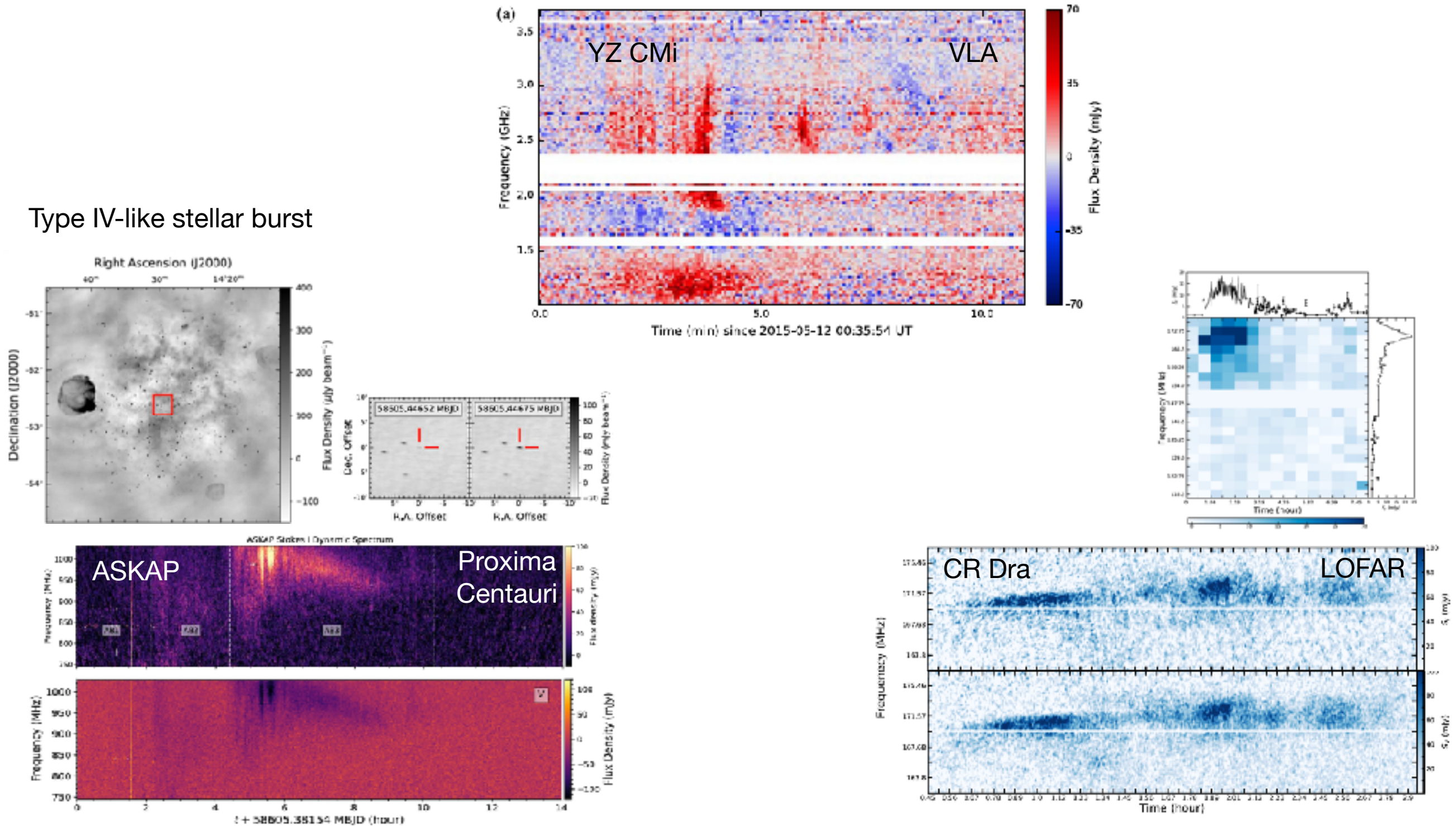
$$\sum_{\text{all baselines}} V(u, v) = \text{dynamic spectrum at phase center}$$

Stars

- Stellar dynamic spectra → bursts

[Villadsen & Hallinan, 2019; Zic et al., 2020; Callingham et al., 2021]

Type IV-like stellar burst



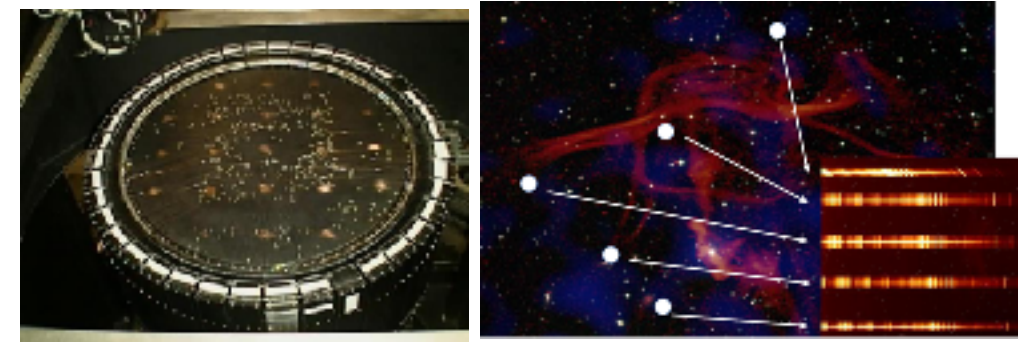
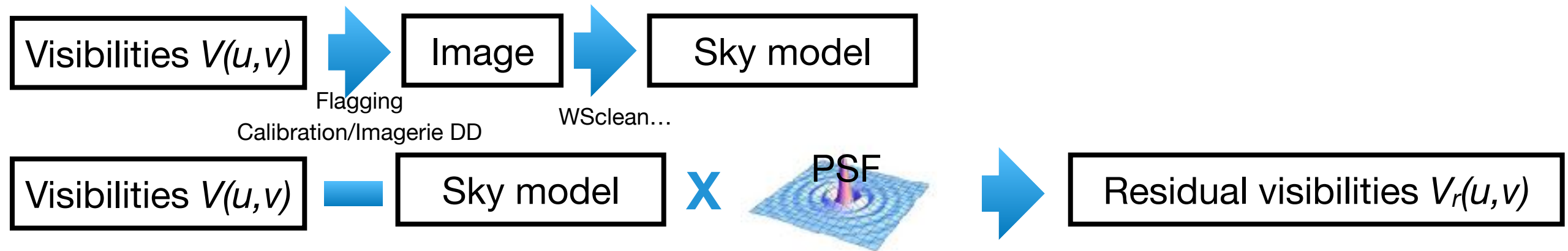
Stars, Exoplanets

$$\Sigma V(u,v) \cdot \vec{e}^{ik \cdot b} = \text{dynamic spectrum in direction } \vec{k}$$

→ digital Integral Field Spectroscopy

[Tasse et al., Nat. Ast., in revision]

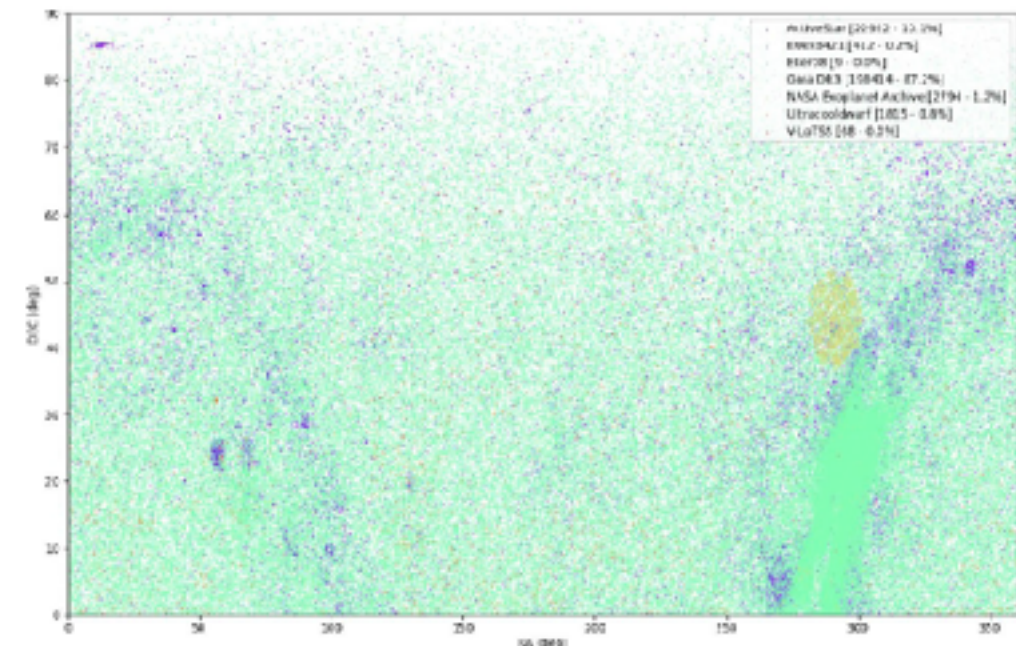
Improvement :



$$\Sigma V_r(u,v) \cdot \vec{e}^{ik \cdot b} = \text{dynamic spectrum of the fluctuations in direction } \vec{k}$$

→ DynSpecMS :

hundred directions per field possible
efficiency +++
input catalog (stars & exoplanets)



Stars, Exoplanets



- First detection of bursts in "high-variability" LOFAR dynamic spectra
- Confirmation via re-imaging
- Consistent with Star-Planet Interaction emission
- Application to NenuFAR

[Tasse et al., Nat. Ast., in revision;
Zhang et al., A&A, in revision]

- First stellar type II burst ?

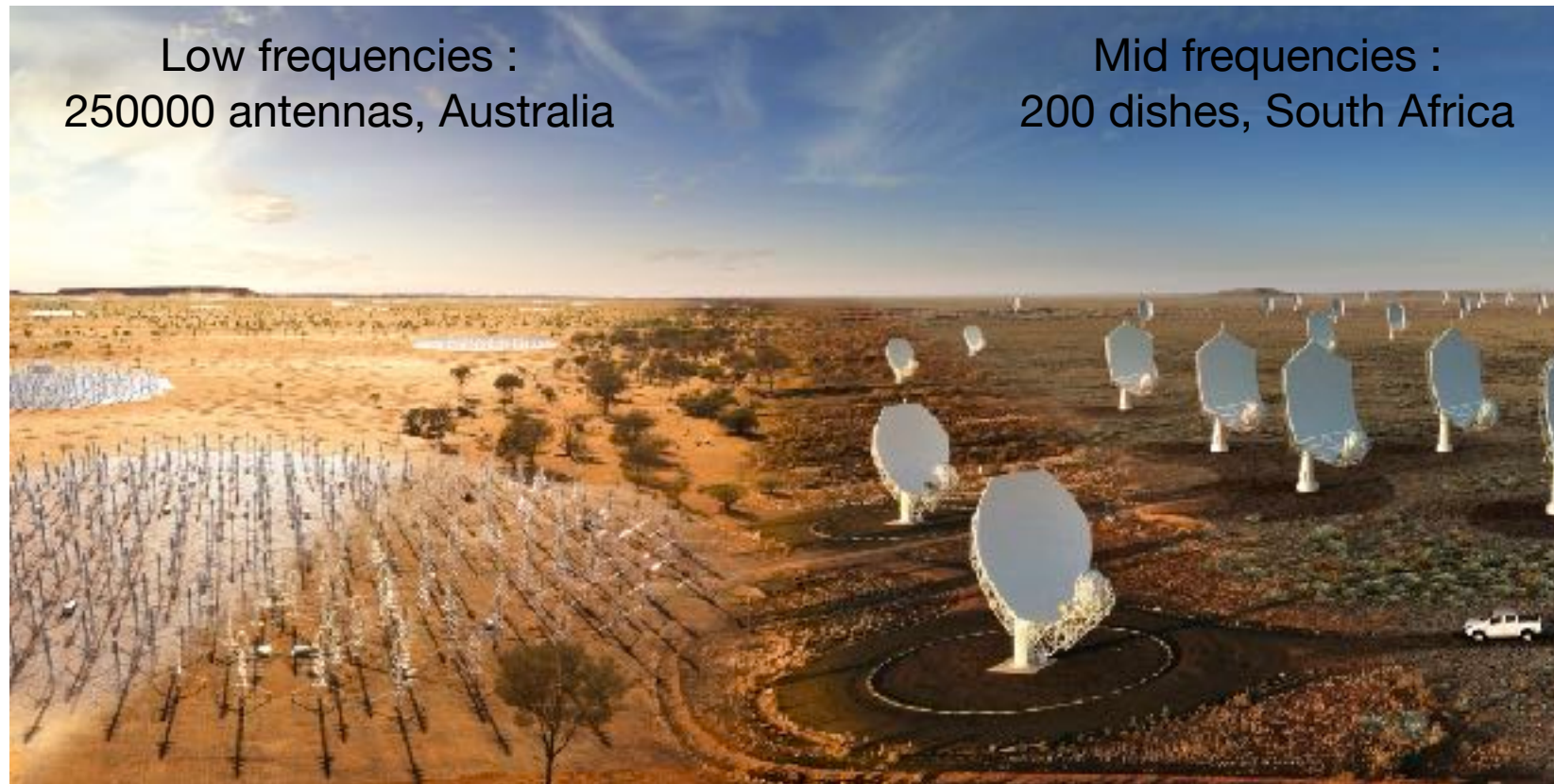
[Tasse et al., Nat. Ast., in revision; Callingham et al., Nature, in revision]

- Heroic times
 - Jupiter, Sun
- Single dishes and phased arrays: t-f studies
 - Jupiter (<1990, 1990-2000, 2000-2010, 2010-2020, 2020+)
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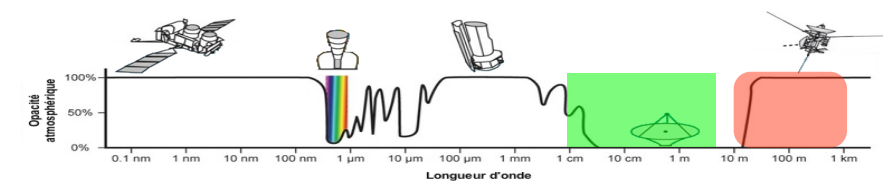


Low frequencies :
250000 antennas, Australia

Mid frequencies :
200 dishes, South Africa



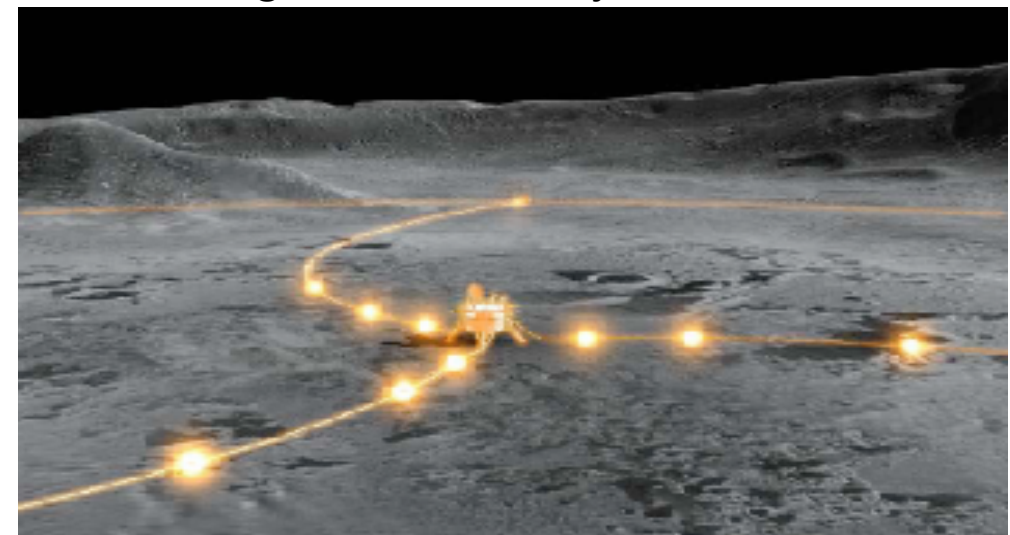
• Problematics:



VLF radio interferometry from space



Large antenna array on the Moon



Thank you.

Questions ?